SYLLABUS



KARPAGAM ACADEMY OF HIGHER EDUCATION (Deemed to be University) (Established Under Section 3 of UGC Act 1956) COIMBATORE-21

(For the candidates admitted from 2018 onwards)

DEPARTMENT OF PHYSICS

SUBJECT: THERMODYNAMICS AND STATISTICAL MECHANICS SEMESTER: II SUB.CODE:19PHP201 CLASS: I M.Sc PHYSICS

Instruction Hours / week: L: 4 T: 0 P: 0

Marks: Internal: 40

External: 60 Total: 100 End Semester Exam: 3 Hours

Course Objectives

- Thermodynamics is an important branch of physics, which helps us to understand the different phenomena in the evolution of the universe.
- This paper gives a basic idea about the laws of thermodynamics and statistical processes. Course Outcomes (COs)
 - 1. Identify and describe the statistical nature of concepts and laws in thermodynamics, in

particular: entropy, temperature, chemical potential, Free energies, partition functions.

2. Use the statistical physics methods, such as Boltzmann distribution, Fermi-Dirac and

Bose-Einstein distributions to solve problems in physical systems.

3. Apply the concepts and laws of thermodynamics to solve problems in thermodynamic

systems such as gases, heat engines and refrigerators etc.

UNITI- LAWS OF THERMODYNAMICS

Some consequences of the laws of thermodynamics – Entropy – Calculation of entropy changes in reversible processes. The principle of increase of entropy – Thermodynamic potentials – Ehthalpy, Helmholtz and the Gibbs functions – Phase transitions – The Clausius-Clayperon equation – Van der Waals equation of state.

UNIT II- KINETIC THEORY

Distribution function and its evolution – Boltzmann transport equation and its validity – Boltzmann's H-theorem – Maxwell-Boltzmann distribution – Transport phenomena – Mean free path- Conservation laws – Hydrodynamics (No derivation).

UNIT III- CLASSICAL STATISTICAL MECHANICS

Maxwell Boltzmann distribution law: Evaluation of constants - Maxwell's law of distribution of velocities - Most probable speed, Average speed, Root mean square speed - Principle of equipartition of energy - Partition function - Condition for applicability of M.B statistics - Non

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degenerate and degenerate systems - Maxwell velocity distribution in a given direction - Total internal energy of an ideal gas - Molar heat capacity of a gas at constant volume – Entropy - Helmholtz free energy - Pressure and equation of state of an ideal gas - Limitation of M.B method.

UNIT IV- QUANTUM STATISTICAL MECHANICS

B.E energy distribution for energies in the range E to E + dE - Condition for B.E distribution to approach classical M.B distribution - Bose temperature - Bose Einstein condensation - Planck's law from B.E law - Fermi Dirac distribution law (no derivation) - FD law for the energies in the range E to E+dE – Fermi energy - Effect of temperature - Energy distribution curve - Free electron in a metal - Fermi temperature and Thermionic emission - Richardson Dushmann Equation - Comparison of MB,BE and FD statistics.

UNIT V- APPLICATIONS OF QUANTUM STATISTICAL MECHANICS

Ideal Bose gas : Photons – Black body and Planck radiation – Photons – Specific heat of solids – Liquid Helium.

Ideal Fermi gas : Properties – Degeneracy – Electron gas – Pauli paramagnetism Ferromagnetism : Ising and Heisenberg models.

SUGGESTED READINGS:

- 1. Agarwal B.K. and M. Eisner, 3rd edition, 2013, Statistical Mechanics, New age international Limited, New Delhi.
- 2. Reif F., 2008, Fundamentals of Statistical and Thermal Physics, (Reprint), McGraw Hill International Edition, Singapore.
- 3. Gupta and Kumar, reprint, 2014, Elements of Statistical Mechanics, Pragati Prakashan, Meerut.
- 4. Huang K., 2nd edition, 2014, Statistical Mechanics, Wiley Eastern Limited, New Delhi
- 5. Sears N. and L. Salinger, 2013, Thermodynamics, 3rd Ed., Narosa Publishing House, New Delhi.
- 6. Greiner W., L. Neise and H. Stocker, 1st edition, 2007, Thermodynamics and Statistical Mechanics, Springer Verlag, New York.
- 7. Singh. K. and S.P. Singh reprint 2016, Elements of Statistical Mechanics, S. Chand & Company Ltd., New Delhi.



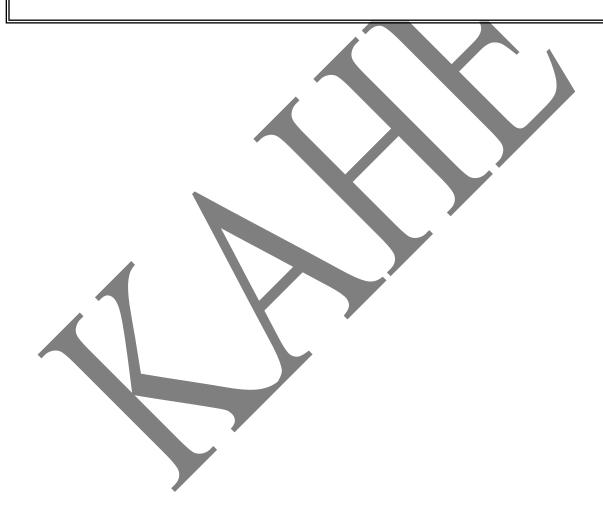
CLASS: I M.Sc Physics

COURSE CODE: 19PHP201

COURSE NAME: Thermodynamics And Statistical Mechanics UNIT: I BATCH-2019-2021

UNIT-I

Laws of Thermodynamics: Some consequences of the laws of thermodynamics - Entropy -Calculation of entropy changes in reversible processes. The principle of increase of entropy -Thermodynamic potentials - Ehthalpy, Helmholtz and the Gibbs functions - Phase transitions -The Clausius-Clayperon equation – Van der Waals equation of state.





THERMODYNAMICS

Thermodynamics is the science of studying the changes that occur within a system in relation to its interaction with its surroundings according to a series of laws formulated that are considered valid for all systems. Thermodynamics allows scientists to study the potential reactions and interactions of systems that exist only in theory, or be such that they cannot be recreated or contained in a laboratory for study.

The empirical facts of thermodynamics are comprehended in its four laws. The first law specifies that energy can be exchanged between physical systems as heat and thermodynamic work. The second law concerns a quantity called entropy, expresses limitations, arising from what is known as irreversibility, on the amount of thermodynamic work that can be delivered to an external system by a thermodynamic process.

THERMODYNAMIC SYSTEMS

An important concept in thermodynamics is the "system". A physical system is the region of the universe under study. A system is separated from the remainder of the universe by a boundary which may be imaginary or not, but which by convention delimits a finite region. The possible exchanges of work, heat, or matter between the system and the surroundings take place across this boundary. There are five dominant classes of systems:

- 1. Isolated Systems matter and energy may not cross the boundary.
- 2. Adiabatic Systems heat may not cross the boundary.
- 3. Diathermic Systems heat may cross boundary.
- 4. Closed Systems matter may not cross the boundary.

Open Systems – heat, work, and matter may cross the boundary.

LAWS OF THERMODYNAMICS

The four main laws of Thermodynamics are,

Zeroth Law - if two systems each are in equilibrium with a third system, then they must are also be in thermal equilibrium with each other.

First Law - if heat is added to a system, some of that energy stays in the system and some leaves



the system.

Second Law - no reaction is 100% efficient and all energy wants to flow and spread to areas with less energy.

Third Law - it is impossible to cool an object to absolute zero because all processes will cease before absolute zero is reached, this is commonly called the state of entropy.

ENTROPY:

Entropy is a defined function of the thermal state of a body and is not affected in any way by the manner in which a particular state is reached. The change in entropy passing from one state A to another state B is given by $S_B - S_A = \int_A^B \frac{dQ}{T}$

where dQ is the quantity of heat absorbed or rejected at a temperature T in going from state A to state B.

- (i) Entropy of a system remains constant during an adiabatic change,
- (ii) Entropy of a system remains constant in all reversible processes.
- (iii) Entropy of a system increases in all irreversible processes.

CALCULATION OF ENTROPY

(i) Entropy of an ideal gas

Consider n gram molecules of an ideal gas occupying a volume V at a pressure P and temperature T. Let quantity of heat dQ be given to the gas, then I law of thermodynamics is dQ=dU+dW. If C_v is the heat capacity of gas at constant volume, dT is rise in temperature and dV represents change in volume.

Then,
$$dU = C_v dT$$
 and $dW = pdV$

$$\implies$$
 dQ = C_v+pdV

From second law of thermodynamics, the change in entropy is

$$dS = dQ/T = (C_v dT)/T + pdV / T$$
 ------(1)

If select some arbitrary state1 at temperature T_0 , pressure P_0 and volume v_0 in which the entropy of the gas is change in entropy during state1 to state2 at temperature T, pressure P and volume V is given by,

$$\Delta s = s - s_0 = \int_1^2 \frac{C_{\nu} dT}{T} + \int_1^2 \frac{P \, dV}{T} \quad ------(2)$$

(a) Value of S terms of temperature and volume

From the equation of state of an ideal gas

$$\implies$$
 P =nRT /

Sub. value of P in eqn (2).

$$\Delta S = \int_{1}^{2} \frac{C_{v} dT}{T} + nR \int_{1}^{2} \frac{dV.T}{V.T}$$

If Cv be assumed to constant, equ.

$$\Delta S = Cv \log_e T/T_0 + nR \log_e V/V_0 \quad \text{(state 1 and 2)} \quad \text{(4)}$$

For a case of isothermal expansion, T =

 \therefore The change in entropy of gas in the case becomes

$$\Delta S = nR \log_{V_0} V \tag{5}$$

(b) Value of S in terms of temperature and pressure:

PV = nRT

$$V = nRT/P$$

and PdV + VdP = nRdT ------ (5a)

So that
$$PdV = nRdT - VdP = nRdT - \frac{nRTdP}{P}$$
 using (5a)

Sub. the value PdV in eq. (2)

$$\Delta S = \int_{1}^{2} Cv \, \frac{dT}{T} + \, nR \int_{1}^{2} \left(\frac{dT}{T} - \frac{dP}{P}\right)$$



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$$= C_{V} \log_{e} \frac{T}{To} + nR (\log_{e} \frac{T}{To} - log_{e} \frac{P}{Po})$$

For an ideal gas, $C_{P} - C_{V} = nR$
$$\Delta S = C_{P} \log_{e} \frac{T}{To} + (C_{P} - C_{V}) log_{e} \frac{P}{Po}$$

For isothermal change $T=T_0$, therefore change in entropy of the gas,

(ii) Entropy of steam

Let consider mass m of ice at absolute temp. T_1 , find the total gain in entropy when ice changes into steam at absolute temperature T_2 .

If small amount of heat dQ is given to a substance at temperature T, the change in entropy is dS = dQ/T.

To convert mass m of ice at T_1K into water at same temperature, the amount of heat required = mL_i , where L_i is the latent heat of ice.

Change in entropy during this process = mL_i/T ------ (i)

When mass m of water at T_1 K is heated to T_2 K, the change in entropy,



$$\Delta S = \int_{T_1}^{T_2} \frac{dQ}{T} = \int_{T_1}^{T_2} \frac{mCdT}{T} = mC \log_c T_2 / T_1$$
 ------ (ii)

To convert mass m of water at T_2C into steam at same temperature, the amount of heat required = mL_s, where L_s is latent heat of steam.

Change in entropy during this process = mL_s/T_2 ------ (iii)

Total gain in entropy = $\frac{mLi}{T_1} + mCloge\left(\frac{T_2}{T_1}\right) + mL_s/T_2$

PRINCIPLE OF INCREASE OF ENERGY OR DEGRADATION OF ENERGY

The entropy of a system remains constant in reversible cyclic process but increased inevitably in all irreversible process. Since a reversible process represents a limiting ideal case, all actual process are inherently irreversible. It means that as cycle after cycle of operation is performed, the entropy of the system increase and tends to a maximum value. This is the principal of increase of entropy and may be stated as " The entropy of an isolated or self contained system either increase or remains constant according as the process it undergoes are irreversible or reversible". Analytically it may be expressed as greater than 0; where the equality sign refers to reversible processes and the inequality sign to irreversible processes. Therefore the necessary and sufficient conditions of equal brim of a self contained system is that it's entropy should be maximum and it cannot be greater than zero.

Since all physical operation in the universe are irreversible for every such operations performed, a certain amount of energy become unavailable for useful work and is added to the universe in the form of heat through friction, conduction or radiation. In this way in a distant future on account of irreversibly all energies existing in different forms will be converted into heat energy and will not be available for conversion into mechanical work i.e." The available energy of the universe is tending toward zero " it will correspond to a state of maximum entropy and all temperature difference between various bodies of the universe will be equalized due to convection etc. No heat engine will then be able to work in this state because no heat flow would be possible due to the uniformity of the temperature throughout the universe .This is called the principal of degradation of energy is conserved it is transformed into a form which is unavailable



for work . Thus the energy is "running down hill" and the universe is marching toward stage of die a "head - death".

With an increase in entropy, the thermal agitation and hence disorder of molecules of a substance increase, i.e increase of entropy implies a transition from order to discord. Thus the principal of increase of entropy is intimately connected with the less ordered state of affairs. According to it, a system posing high entropy should be in great disorder or chaos. Thus the entropy of a substance in gaseous state is more than in liquid state, because the molecules are free to move about in great disorder in a gas than in a liquid. Moreover the entropy is more in a liquid state than in the solid state, as the molecules are more free to move in a liquid than in a solid. Hence when ice is converted into water and then into steam, the entropy and disorder of molecules increase. On the other hand when the steam is converted into water and then in to ice, the entropy and disorder of molecules continually decreased. Thus when the temperature of a system is decreased, the amount of entropy and disorder in it decreased. Entropy of the substance is therefore said to be a measure of the degree of disorder prevailing among it's molecules just as the temperature is a measure of the degree of hotness of a substance at the absolute zero of temperature the thermal motion completely disappears so that the disorder and hence the entropy tends to zero and the molecules of a substance are in perfect order i.e well arranged.

By summarizing the above arguments, say that the entropy of any isolated system increase and approaches more or less rapidly to the inert state of maximum entropy. We may recognize this fundamental law of physics to be an inherent tendency of nature to be processed from a more ordered state to a less ordered one or from a less disordered to a more disordered state or other words that the ultimate destiny of universe is not order but chaos.

THERMODYNAMIC POTENTIALS:

The thermodynamics variables such as pressure P, Volume V, temperature T and entropy S, define the stole of thermo dynamical system. A relation b/n them exists because of the two thermodynamic laws.

dQ = dU + PdV

dQ = TdS



Combining

TdS = dU + PdVdU = TdS - PdV

Any two of above variables are independent and with the help of above relation remaining variables be determined. These relations are termed as thermodynamic potentials of thermodynamic fn.

Enthalpy or Heat content H

Thermodynamical phenomena at constant pressure are expressed in terms of another function called enthalpy or heat content of the system.

--- (1)

$$H = U + PV$$

diff. dH=dU+PdV+VdP

$$=$$
 (TdS-PdV) + PdV+VdP

= TdS + VdP

dH = Tds

because process is carried at constant pressure. Since Tds = dQ, we find dH=dQ enthalpy represents the quantity of heat given to the system from an external source & hence the name heat content.

Let H_i and H_f be the initial and final enthalpy,

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H_f-H_i = Q
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The change in enthalpy during an isobaric process equal too the heat transferred.

H has an important property in porous plug exp let P_i and V_i be the initial pressure & volume of a gas before passing through porous plug. Similarly pf and vf be the similar quantities of the gas after passing thro' the porous plug.

External work done by $gas = P_f V_f - P_i V_i$

This work done at the cost of internal energy of the because no heat exchanges b/n gas and surrounding, suppose Ui and Uf be the initial and final internal energy.

 $U_i - U_f = P_f V_f - P_i V_i$

 $\frac{\partial}{\partial S}\left(\frac{\partial H}{\partial P}\right) = \frac{\partial}{\partial P}\left(\frac{\partial H}{\partial S}\right)$

 $\frac{\partial T}{\partial P}$)s

 $\left(\frac{\partial V}{\partial P}\right)_{V}$

$$U_i + P_i V_i = P_f V_f + U_f$$
(or) $H_i = H_f$

(3)

(4)

Thus in throttling process, the initial and final enthalpy remain same.

Taking partial diff. of H w.r. to independent variables S and P,

$$\implies \quad (\frac{\partial H}{\partial P})_{S} = V \text{ and } (\frac{\partial H}{\partial S})_{P} = T$$

As dH is perfect diff.

Using (3), we get

Which is third thermodynamical relation.

Helmholtz Function F

On combining I & II law of thermodynamics,

dU = Tds - dW

suppose the temp of the system remain constant, then

d(TS) = Tds

$$\implies dU = d(TS) - dW$$



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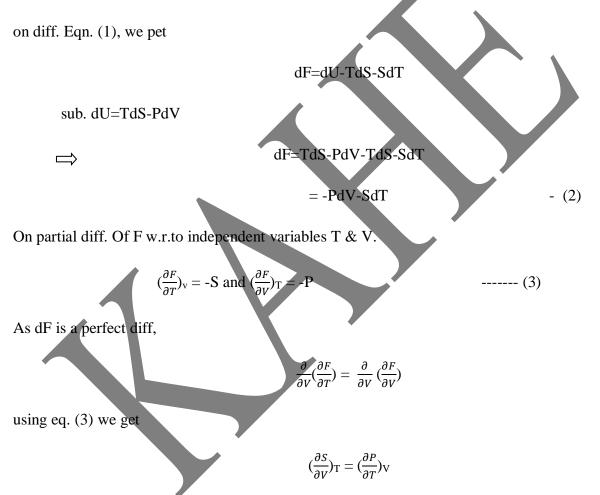
d(U-TS) = -dW

where the fn. F=U-TS

$$\implies$$
 dF=-dW

----- (1)

is called Helmholtz fn. Or Helmholtz free energy, which represents that in revisable isotheral process, the work done by the system is equal to decrease in Helmholtz Fn. F is also called as work fn.



This eqn. given a relation b/n 4 thermo dynamical variable P,V,S and T. This is second thermodynamical relation

Gibbs potential (G)

If thermodynamic process is isothermal and isobaric (dp=0) then from qn.(2) we get



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-- (6)

(7)

dH=T(ds)

= d (TS) [from Helmholtz fn.]

d(H-TS) = 0

dG = 0

Where G = H - TS

G = U + PV - TS

is called gibb's fn. or free energy. On diff. (5), we get

dG = dU + PdV + VdP - TdS - SdT

 $= (T dS - p d\psi) + p d\psi + V dP - T dS - S dT$

dG = VdP-SdT

Talking partial derivatives of a w.r. to independent variable P and T, we get

 $\left(\frac{\partial G}{\partial P}\right)_{\mathrm{T}} = \mathbf{V} \text{ and } \left(\frac{\partial G}{\partial T}\right)_{\mathrm{P}} =$

As dG is perfect diff ., $\frac{\partial}{\partial T} \left(\frac{\partial G}{\partial P} \right) = \frac{\partial}{\partial P} \left(\frac{\partial}{\partial P} \right)$

This is called fourth thermodynamical relation.

PHASE TRANSITIONS

Simple substances are capable of existing in phases of three types: solid, liquid and gas. The three lines, in phase diagram separating these planes are called phase equilibrium lines. The common point A where three lines meet is called *triple point*; at this unique temperature and pressure all three phases can coexist in equilibrium with each other. Point C is the critical point at which liquid gas equilibrium line ends. The volume change ΔV between liquid and gas then approached zero; beyond C there is no further phase transition since only one fluid phase exist.



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CLASUSIS CLAYPERON EQUATION

Maxwell second thermo dynamical relation is

$$\left(\frac{\partial S}{\partial V}\right)_{\rm P} = -\left(\frac{\partial P}{\partial T}\right)_{\rm V}$$

Multiply both sides by T,

 $T(\frac{\partial S}{\partial V})_{T} = T(\frac{\partial P}{\partial T})_{V}$ TdS = dQ $(\frac{\partial Q}{\partial V})_{T} = T(\frac{\partial P}{\partial T})_{V}$

From II law of thermodynamics,

 $(\frac{\partial Q}{\partial v})_T$ represent the quantity of heat absorbed or liberated per unit charge in volume at constant temp. This means that at constant temp. The heat absorbed or liberated bring out simply a change in the volume of the substance. Therefore this amount of heat absorbed or liberated at constant temp must be the latent heat and change in volume must be due to change of state. Considering a unit mass of the substance let L be the latent heat when the substance change in volume from V₁ to V₂ at constant temp. then,

$$\delta Q = L \text{ and } \delta V = V_2 - V$$

$$\left(\frac{L}{V_2 - V_1}\right) = T \left(\frac{\partial P}{\partial T}\right)_V$$

$$\frac{L}{V_2 - V_1} = T\frac{\partial P}{\partial T}$$

$$dP/dT = \frac{L}{T(V_2 - V_1)}$$

which is called Clausius Clapyeron latest heat equation.



VAN DER WAAL'S EQUATION OF STATE

Consider the volume occupied by the gas molecules negligible compared with the total volume of gas and the molecules exert no appreciable forces on one another. It is evident that both these assumptions cannot be exactly true for actual gases particularly at high pressure. In driving van der waals eqn. Of state the effect both these factor is taken into account.

Due to the finite size of molecules, the free space available for their movement is less than the actual measured volume of the gas. Also the number of collisions with the walls of containing vessel, and the pressure will be greater than the calculated by simple theory. The actual volume can be brought about by subtracting a career term b from the measured volume and using (v-b) in place of V in ideal gas equation.

Let XY be the portion of boundary wall. Consider a molecule A in the interior far from the boundary wall. It is surrounded by other molecules equally distributed in all directions. Those molecules exert attractive force on molecule A, when averaged out, over a sufficient interval of time they cancel out and net cohesion force will be zero. On the other hand the molecule B is as rear the boundary as it can go. In this case the molecular distribution is only along one side. The adhesive force between the gas molecules and the boundary walls are always must smaller than the cohesive force, between the gas molecules. The force on B due to each adjacent molecule can be resolved into components to the boundary wall. The parallel components cancel out on the average but the perpendicular components will result a field of force acting inwards on the molecules near the boundary wall. Thus whenever a molecule will strike the walls of the containing vessel at B to contribute its share towards the total gas pressure, the measured pressure P is loss than the ideal pressure calculate on the assumption that the cohesive force is P, the add a correction term P, to the measured pressure P and use $(P+P_1)$ in place of P in ideal gas.

On using both corrections in ideal gas equation, we get for a gram molecule of a gas.

$$(P+P_1)(V-b) = RT$$
 (1)

The value of P_1 is to the number of molecules striking in area of the wall in unit time & to the intensity of the field of force. Both of these factors are proportional to the density of the gas.



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 $p_1 = a\rho^2$

C->constant

 $\rho \alpha 1/V$

Hence $p_1 = a/V^2$, where a is constant.

Sub. this value of p_1 , in equation (1),

 $(p+a/V^2)(v-b)=RT$

This is van der waal's equation of state. This is the simplest and the most well known equation of stove for real gas.

Another useful form of the equation of state of a real gas is

PV=A+B/V+C/V \rightarrow (3)

A,B,C,... are from of temp and are called virial coeff.

For an ideal gas it is evident that A=RT and all other viral coeff, are zero.

Van der walls equation can be but in virial form as

Equation (2) rewritten as,

(or)

1=RT/((P+a/V²)(v-b))

 $PV=RT(l-a/PV^2)^{-1} (l-b/v)^{-1}$ \rightarrow (4)

The correction terms $a/PV^2 \& b/V$ are both small composed with unity provided the gas is not too much compressed. Using binomial theorem & neglecting the terms of higher power a l/V, equation V1, becomes,

 $PV=RT(1-A/PV^2)(1+B/V)+b^2/v^2$

Since PV=RT approx,



PV=RT+(RTb-a)/v+RTb²/v² →(5)

This is van der waal's equation in virial form having only three virial coeff., A,B,&C

A=RT, B=RTb-a, C=RTb²



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Possible Questions

1. Define entropy.

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- 2. State first and second law of thermodynamics.
- 3. What are called thermodynamic potentials?
- 4. Define the term drift variation and collision interaction.
- 5. Write down the Clausius-Clapeyron equation for latent heat.
- 6. Distinguish between real and ideal gas.
- 7. State third law of thermodynamics.
- 8. Define Helmholtz function.
- 9. Calculate the entropy change in reversible processes.
- 10. Derive Vander Waals equation of state.
- 11. Derive Clausius-Clapeyron equation.
- 12. Obtain Helmholtz free energy Gibbs function.
- 13. Define the term "phase transition".
- 14. Obtain three different thermodynamic potentials.
- 15. Explain principle of increase of entropy.

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(For the candidates admitted from 2018 onwards)

DEPARTMENT OF PHYSICS

UNIT I (Objective Type/Multiple choice Questions each Questions carry one Mark)

THERMODYNAMICS AND STATISTICAL MECHANICS

	TAKT-A(Oninie Examination)							
S.No.	QUESTIONS	OPTION 1	OPTION 2	OPTION 3	OPTION 4	ANSWER		
	The term "thermodynamics" comes from							
	Greek. words "therme" and "dynamis" which							
1	means	Heat power	Heat transfer	Heat energy	Heat motion	Heat power		
	The term "thermodynamics" was first used in	Rudolph	William					
2	1849 in the publication of a	Clausius	Rankine	Lord Kelvin	Thomas Savery	Lord Kelvin		
	The macroscopic approach to the study of							
	thermodynamics does not require a							
	knowledge of the behavior of individual	Dynamic	Static	Statistical	Classical	Classical		
3	particles is called	thermodynamics	thermodynamics	thermodynamics	thermodynamics	thermodynamics		
	What law asserts that energy is a	First law of	Second law of	Third law of	Zeroth law of	First law of		
4	thermodynamic property?	Thermodynamics	Thermodynamics	Thermodynamics	Thermodynamics	Thermodynamics		
	What law asserts that energy has quality as	First law of	Second law of	Third law of	Zeroth law of	Second law of		
5	well as quantity?	Thermodynamics	Thermodynamics	Thermodynamics	Thermodynamics	Thermodynamics		
					The entropy-			
	The first law of thermodynamics is based on		Conservation of	Conservation of	temperature	Conservation of		
6	which of the following principles?	mass of energy	energy	MOMENTUM	relationship	energy		
		microscopic	macroscopic	homogeneous	heterogeneous	macroscopic		
7	Thermodynamics is applicable to	systems only	systems only	systems only	systems only.	systems only		
8	Which is not true about thermodynamics ?	it ignores the	it involves the	it is concerned	it is not	it is not applicable		

PART –A(Online Examination)



		internal structure of atoms and molecules	matter in bulk	only with the initial and final states of the system	applicable to macroscopic systems.	to macroscopic systems.
	A system that can transfer neither matter nor		an isolated		a homogeneous	
9	energy to and from its surroundings is called	closed system	system	an open system	system	an isolated system
	Which of the following is incorrect, for an					
10	ideal gas ?	PV = nRT	V = nRT/P	P=nRT/V	P =RT	PV = nRT
	The heat capacity at constant pressure is related to heat capacity at constant volume by					
11	the relation	Cp-R =Cv	Cv-R =Cp	Cp-Cv =R	R-Cp =Cv	Cp-Cv =R
	A system is in equilibrium if the temperature is the same throughout the entire					
12	system.	Static	Thermal	Mechanical	Phase	Thermal
	A system is in equilibrium if there is no change in pressure at any point of the					
13	system with time.	Pressure	Thermal	Mechanical	Phase	Mechanical
14	If a system involves two phases, it is in equilibrium when the mass of each phase reaches an equilibrium level and stays there.	Chemical	Thermal	Mechanical	phase	Phase
	A system is in equilibrium of its chemical composition does not change with					
15	time, i.e., no chemical reaction occurs.	Chemical	Thermal	Mechanical	Phase	Chemical
	A system is said to be in thermodynamic equilibrium if it maintains	Mechanical and	Thermal and	Thermal, mechanical and	Thermal, phase, mechanical and	Thermal, phase, mechanical and
16	equilibrium.	phase	chemical	chemical	chemical	chemical
	What is a process with identical end states				Either path or	
17	called?	Cycle	Path	Phase	phase	Cycle
	What is a process during which the		Isothermal		Isometric	Isothermal
18	temperature remains constant?	Isobaric process	process	Isochoric process	process	process
	What is a process during which the pressure		Isothermal		Isometric	
19	remains constant?	Isobaric process	process	Isochoric process	process	Isobaric process

	What is a process during which the specific		Isothermal	Isochoric or	Isovolumetric	Isochoric or
20	volume remains constant?	Isobaric process	process	isometric process	process	isometric process
	What states that if two bodies are in thermal					
	equilibrium with a third body, they are also in	Zeroth law of	First law of	Second law of	Third law of	Zeroth law of
21	equilibrium with each other?	thermodynamics	thermodynamics	thermodynamics	thermodynamics	thermodynamics
	What is the study of energy and its					
22	transformations?	Thermostatics	Thermophysics	Thermochemistry	Thermodynamics	Thermodynamics
	What is considered as the heat content of a					
23	system?	Enthalpy	Entropy	Internal heat	Molar heat	Enthalpy
	What refers to the amount of heat needed to					
	raise the temperature of an object by one					
24	degree Celsius or 1K?	Heat capacity	Specific heat	Latent heat	Molar heat	Heat capacity
	What is the heat capacity of one mole of					
25	substance?	Molecular heat	Specific heat	Latent heat	Molar heat	Specific heat
	What refers to the measure of the disorder					
26	present in a given substance or system?	Enthalpy	Entropy	Heat capacity	Molar heat	Entropy
			Joule-			
27	Entropy is measured in	Joule/Kelvin	Meter/Kelvin	Meter/Kelvin	Newton/Kelvin	Joule/Kelvin
	What is the energy absorbed during chemical				Enthalpy of	
28	reaction under constant volume conditions?	Entropy	Ion exchange	Enthalpy	reaction	Enthalpy
	Which of the following equation is used to					
	calculate the heats of reaction when ΔG at	Gibbs Helmholtz	Clapeyron	Kirchoffs		Gibbs Helmholtz
29	two temperatutes are given?	equatioin	equation	equation	Nernst equation	equatioin
20	is applicable to macroscopic		.1 1		thermochemical	.1 1 .
30	systems only.	thermochemistry	thermokinetics	thermodynamics	studies.	thermodynamics
24		first law of	second law of	41.2 1	third law of	first law of
31	$\Delta E = q$ -w for an isochoric process	thermodynamics	thermodynamics	zeroth's law	thermodynamics	thermodynamics
32	Who proposed the Carnot cycle?	Sammy Carnot	Sonny Carnot	Sadi Carnot	Suri Carnot	Sadi Carnot
33	Entropy is transferred by	Work	Heat	Energy	Work and heat	Heat
34	Gibb's function is expressed as,	G = H + TS	G = H / TS	G=H-TS	G = H * TS	G=H-TS
		Directly	Directly	Independent of	Inversely	Directly
35	Average kinetic energy of molecules is	proportional to	proportional to	absolute	proportional to	proportional to

		square root of	absolute	temperature	absolute	absolute
		temperature	temperature		temperature	temperature
	The specific heat of a gas in isothermal			Remains		
36	process is	Zero	Negative	constant	Infinite	Infinite
		Less than	Equal to external	More then	Twice the	Twice the external
		external latent	latent heat of	external latent	external latent	latent heat of
37	Latent heat of ice is	heat of fusion	fusion	heat of fusion	heat of fusion	fusion
	The difference between the principal specific					
	heats of nitrogen is 300 J/kg °K and ratio of					
38	the two specific heats is 1.4. then the CP is	1050 J/kg °K	650 J/kg °K	750 J/kg °K	150 J/kg °K	650 J/kg °K
	The mean kinetic energy of one gram-mole					
39	of a perfect gas at absolute temperature T is	1/2 KT	1/2 RT	3/2 KT	3/2 RT	3/2 RT
	The specific heat of a substance at its boiling				Lies between 0	
40	point or melting point	Is zero	Is infinity	Is negative	and 1	Is infinity
	Which of the following variables controls the					
41	physical properties of a perfect gas?	Pressure	Temperature	Volume	Atomic mass	Atomic mass
	A system in which state variables have					
	constant values throughout the system is			isothermal		
42	called in a state of	equilibrium	non- equilibrium	equilibrium	none of these.	equilibrium
	In an adiabatic process can flow in to					
43	or out of the system.	no heat	heat	matter	no matter	no heat
	The mathematical relation for the first law of		$\Delta E = 0$ for a	$\Delta E = -q$ for an		
44	thermodynamics is	$\Delta E = q + w$	cyclic process	isochoric process	$\Delta E = W-q.$	$\Delta E = q + w$
	For an adiabatic process according to first					
45	law of thermodynamics,	$\Delta E = -w$	$\Delta E = w$	$\Delta E = q-w$	$\Delta q = E - w$	$\Delta E = -w$
	The enthalpy change, ΔH of a process is					
46	given by the relation	$\Delta H = \Delta E + p \Delta v$	$\Delta H = \Delta E + \Delta nRT$	$\Delta H = \Delta E + w$	$\Delta H = \Delta E - \Delta n R T$	$\Delta H = \Delta E + p \Delta v$
	The amount of heat required to raise the					
	temperature of one mole of the substance by		molar heat			molar heat
47	1 K is called	heat capacity	capacity	molar heat	molar capacity.	capacity
48	Which of the following is not correct ?	H=E+PV	H-E=PV	H-E-PV=0	H=E-PV	H=E-PV
	The enthalpy of a system is defined by the					
49	relation	H=E+PV	H=E-Pv	E=H+PV	PV+E-H	H=E+PV

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(For the candidates admitted from 2018 onwards)

DEPARTMENT OF PHYSICS

UNIT II (Objective Type/Multiple choice Questions each Questions carry one Mark)

THERMODYNAMICS AND STATISTICAL MECHANICS

	raki –A(Onnie Examination)							
S.No.	QUESTIONS	OPTION 1	OPTION 2	OPTION 3	OPTION 4	ANSWER		
1	Which of the following statements is TRUE for an ideal gas, but not for a real gas?	PV = nRT	An increase in temperature causes an increase in the kinetic energy of the gas	The total volume of molecules on a gas is nearly the same as the volume of the gas as a whole	No attractive forces exists between the molecule of a gas	PV = nRT		
2	The molecules of a gas moving through space with some velocity possesses what kind of energy?	Translational energy	Spin energy	Rotational kinetic energy	Sensible energy	Translational energy		
3	Molar specific heat at constant volume is C_v for a monoatomic gas is.	3/2 R	5/2 R	3R	2R	3/2 R		
4	If the pressure in a closed vessel is reduced by drawing out some gas, the mean free path of the molecules.	Is decreased	Remains unchanged	Is increased	Increases or decreases according to the nature of the gas	Is increased		
5	Cooking gas containers are kept in a lorry moving with uniform speed. The temperature of the gas molecules inside	Increase	Remain same	Decrease	Decrease for some, while increase for	Remain same		

PART –A(Online Examination)



	will.	, , , , , , , , , , , , , , , , , , ,	· · · · · · · · · · · · · · · · · · ·		others	
		Temperature				
	1	become four	Temperature	Temperature	Temperature	
	1	times at	becomes two	become one	becomes half at	
	I I	constant	times at constant	fourth at	constant	Temperature become four
6	Volume of gas become four times if.	pressure	pressure	constant pressure	pressure	times at constant pressure
		Inelastic rigid	Perfectly elastic	Perfectly elastic	Inelastic non-	Perfectly elastic rigid
7	Molecules of a gas behave like.	sphere	rigid sphere	non-rigid sphere	rigid sphere	sphere
	At absolute zero temperature, pressure of			One atmospheric		
8	a gas will be	Zero	Po * 273	pressure	Po * 76	Zero
1		Isobaric		Isothermal	Isotonic	
9	Boyle's law holds for an ideal gas during	changes	Isochoric changes	changes	changes	Isothermal changes
			Charle's law and			Charle's law and Boyle's
10	Kinetic theory of gases provide a base for	Charle's law	Boyle's law	Boyle's law	stefans law	law
11	In Boyle's law what remains constant.	PV	TV	V/T	V/T	PV
12	S.I. unit of universal gas constant is	cal/°C	J/molK	J/mol	J/kg	J/molK
	i			Number of		
		Collision on		collisions per		
	At constant volume, temperature is	walls will be	Collisions will be	unit time will	Collisions will	Number of collisions per
13	increased. Then.	less	in straight lines	increase	not change	unit time will increase
		Has only two	Can have any	Has a unique	Depends upon	
	,	values Cp and	value between 0	value at a given	the mass of the	Has only two values
14	The specific heat of a gas	Cv	and ∞	temperature	gas	Cp and Cv
					Real and at	
	1		Perfect and at		constant	
		Perfect and of	constant	Real and of	temperature	
		constant mass	temperature but	constant mass	but variable	Perfect and of constant
15	For Boyle's law to hold the gas should be.	and temperature	variable mass	and temperature	mass	mass and temperature
		At high	At normal	At low		
	Every gas (real gas) behaves as an ideal	temperature and	temperature and	temperature and	_	At high temperature and
	gas.	low pressure	pressure	high pressure	low pressure	low pressure
17	According to kinetic theory of gasses at	Water freezes	Liquid helium	Molecules	Liquid	Molecules motion stops

	absolute zero temperature		freezes	motion stops	hydrogen	
					freezes	
					not equal to	
18	For an ideal gas Cp and Cv is	grater than one	less than one	equal to one	one	less than one
					Not be	
19	An ideal gas is that which can	Be solidified	Liquefied	Not be liquefied	solidified	Not be liquefied
		Directly	Directly		Inversely	
		proportional to	proportional to	Independent of	proportional to	
		square root of	absolute	absolute	absolute	Directly proportional to
20	Average kinetic energy of molecules is	temperature	temperature	temperature	temperature	absolute temperature
		Less than	Equal to external	More then	Twice the	
		external latent	latent heat of	external latent	external latent	More then external latent
21	Latent heat of ice is	heat of fusion	fusion	heat of fusion	heat of fusion	heat of fusion
	The specific heat of a substance at its				Lies between 0	
22	boiling point or melting point	Is zero	Is infinity	Is negative	and 1	Is infinity
	Which of the following properties of gas					
	molecule the one that is same for all ideal					
23	gases at a particular temperature is	Mass	velocity	momentum	kinetic energy	Mass
				Inverse	Inverse	
		Proportional to	Inverse	proportional to	proportional to $T^{1/2}$	
24	Mean kinetic energy of perfect gas is	Т	proportional to T ²	T ⁻²	T ^{1/2}	Proportional to T
	The motion of fluids and the forces acting					
	on solid bodies immersed in fluids and in					
25	motion relative to them is called	dynamics	hydrodynamics	statitics	mechanics	hydrodynamics
	Temperature of a gas can be related to the					
26	motion of the molecules	external	boundary	internal	closed	internal
					1.38 x 10-19	
27	Boltzmann's constant is	1.38 x 10-23 j/k	1.38 x 10-31 j/k	1.38 x 10-32 j/k	j/k	1.38 x 10-23 j/k
28	The word kinetic refers to	locomotion	vibration	motion	resonance	motion

		r				
		closely	not free to	regularly		
29	In gases the particles are	packed	move	packed	far apart	far apart
	8	low density and	high density and	high density but	low density but	- or of or o
30	Gases have	mass	mass	low mass	high mass	low density and mass
		small no of	large no of small	large no of small	large no of	
		small particles	particles in	particles in	large particles	large no of small particles
	what does the Kinetic theory of gases	in constant	constant randam	accelerating	in constant	in constant randam
31	describe?	randam motion	motion	randam motion	randam motion	motion
	which experiment shows how kinetic	1			refration of	
32	theory works?	g by freefall	brownian motion	pin hole camera	light	brownian motion
	what forces are assumed to exist between					
33	particles in a gas	attractive	repulsive	both	no force	no force
34	Kinetic is a / an	latin word	roman word	greek word	arabic word	greek word
	Which one of the following have the		Liquid helium			
35	highest volume?	solid	freezes	gas	Gel	gas
I T		very	very little			
36	Gases are	compressible	compressible	incompressible	not possible	very compressible
		Temperature become four				
		times at				
ļ		constant				
37	The three states of matter depend on	pressure	force	potential energy	biomass	Temperature
38	The term fluids is used for	liquid only	gases only	liquid and gass	gel only	liquid and gas
			6	they have	6	
	Why are liquids and gases termed as		they have	randomly	they are	
39	fluids? Because	they can flow	irregular shape	moving	compressible	they can flow
İ	The Brownian Motion was discovered by			-		
40	the scientist	albert brown	John brown	robert brown	issac brown	John brown
Π	If the car tires are hot, the pressure of gas			same as before	may be high or	
41	molecules in them would be	high	low	heating	low	high
42	Gas can exert	pressure on wall	force on the base	pressure in solid	force in liquid	pressure on wall

	The random motion of smoke or gas					
43	particles in the air is termed as	brueian motion	brownian motion	radom motion	static	brownian motion
			when individual	the total kinetic		
		matter is	particles collide,	energy of	the particles of	when individual particles
		composed of	they undergo no	colliding	matter are in	collide, they undergo no
	All of the following are basic assumptions	very tiny	exchange of	particles remains	continual	exchange of kinetic
44	of the kinetic theory except:	particles	kinetic energy	constant	motion	energy
	For a gas, which pair of variables are					
	inversely proportional to each other (if all					
45	other conditions remain constant)?	P,T	P,V	V,T	n,V	P,T
	The behavior of gases can be fully					
46	determined by	1 law	2 law	3 law	4 law	4 law
		Only one value	Two value of	Three value of	No value of	Only one value of specific
47	Solid and liquids have	of specific heat	specific heat	specific heat	specific heat	heat
		Nominal	Natural	Normal	Normal	
		temperature and	temperature and	temperatuere and	thermodynamic	Normal temperatuere and
48	The term N.T.P stands for	pressurre	pressure	pressure	pressure	pressure
	Boyle's law ie, PV = constant is	All ranges of	Only small range	Steady change of	Atmospheric	Only small range of
49	applicable to gases under	pressures	of pressures	pressures	conditions	pressures
		Nominal	Natural	Normal	Normal	
		temperature and	temperature and	temperature and	thermodynamic	Normal temperature and
50	The term N.T.P stands for	pressure	pressure	pressure	pressure	pressure
	M.B. distribution can be applicable to	identical	indistinguishable			
		molecule	molecule	gas	liquid	identical molecule

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DEPARTMENT OF PHYSICS

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PART –A(Online Examination)



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	If the car tires are hot, the pressure of gas			same as before	may be high or	
41	molecules in them would be	high	low	heating	low	high
42	Gas can exert	pressure on wall	force on the base	pressure in solid	force in liquid	pressure on wall

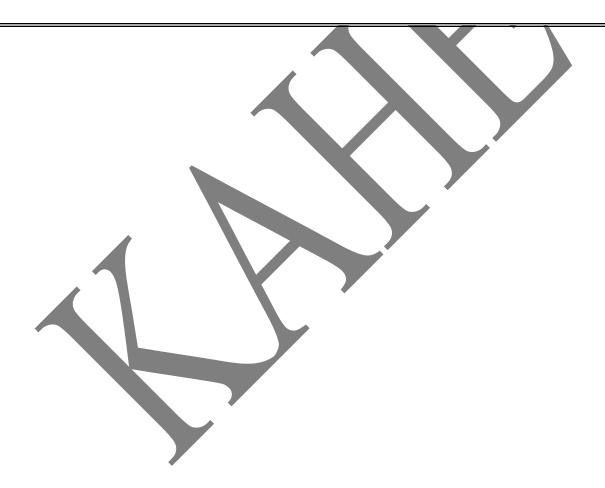
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		when individual	the total kinetic		
	matter is	particles collide,	energy of	the particles of	when individual particles
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determined by	1 law	2 law		4 law	4 law
	Only one value	Two value of	Three value of	No value of	Only one value of specific
Solid and liquids have	of specific heat	specific heat	specific heat	specific heat	heat
	Nominal	Natural	Normal	Normal	
	temperature and	temperature and	temperatuere and	thermodynamic	Normal temperatuere and
	pressurre	pressure	pressure	pressure	pressure
	-		Steady change of	Atmospheric	Only small range of
applicable to gases under	*	of pressures	pressures		pressures
	Nominal	Natural	Normal		
	temperature and	temperature and	temperature and	thermodynamic	Normal temperature and
	pressure	pressure	pressure	pressure	pressure
M.B. distribution can be applicable to		ũ			
	molecule	molecule	gas	liquid	identical molecule
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CLASS: I M.Sc Physics COURSE CODE: 19PHP201 COURSE NAME: Thermodynamics And Statistical Mechanics UNIT: III BATCH-2019-2021

UNIT-III

Classical Statistical Mechanics: Maxwell Boltzmann distribution law: Evaluation of constants -Maxwell's law of distribution of velocities - Most probable speed, Average speed, Root mean square speed - Principle of equipartition of energy - Partition function - Condition for applicability of M.B statistics - Non degenerate and degenerate systems - Maxwell velocity distribution in a given direction - Total internal energy of an ideal gas - Molar heat capacity of a gas at constant volume – Entropy - Helmholtz free energy - Pressure and equation of state of an ideal gas -Limitation of M.B method.





(2)

-----(4)

Maxwell -Boltzmann Momentum Distribution Law for an Ideal Gas

The Maxwell –Boltzmann equation for distribution of energy among the molecules of an ideal gas is

> $n(E)dE = \frac{2\pi N}{(\pi KT)^{3/2}} Ee^{1/2}e^{-E/KT}dE$ -----(1)

All the energy of the gas in the form of KE of its molecules .Therefore

 $E = \frac{1}{2}mv^2 = (\frac{(mv)^2}{2m} = \frac{p^2}{2m})$

Taking differential of this eqn..

$$dE = \frac{p}{m} dp$$

sub..the expression for E and dE in eq(1), the number of molecules n(p)dp whose momentum lie between p and p+dp is

> $n(p)dp = \frac{2\pi N}{(\pi KT)^{3/2}}$ $\frac{1}{2} e^{-p2/2mKT} \left(\frac{p}{m}\right) dp$.p²e^{-p2/2mKT}.dr $\frac{4\pi N}{(2\pi m KT)^{3/2}}$

This eq. is known as Maxwell -Boltzmann law of distribution of momenta among the

molecules of an Ideal Gas.

Evaluation of constant

The total number N of the particles in the system is given by

Where

 $A=e^{-\alpha}$



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From eq. (1)

$$\mathbf{A} = \frac{N}{\sum_{\mathbf{r} \, \mathbf{g}_{\mathbf{r}}} \mathbf{e}^{\frac{-\mathbf{E}\mathbf{r}}{\mathbf{k}\mathbf{T}}}}$$

For continuous variation of energy of free particles of an ideal gas, g_r is replaced by g(E) dE, E_r is replaced by E and the sign of summation is replaced by the sign of integration.

The limits of integration are taken from 0 to ∞ because energy of the particles of an ideal gas is entirely kinetic and so they can have any K.E. The value of g(E) dE for particles with no spin is given by

g(E) dE =
$$2\pi V \left(\frac{2m}{h^2}\right)^{3/2} E^{1/2} dE$$

The integral in eq. (2) is evaluated as follows

$$I = \int_0^\infty g(E) \, e^{-E/kT} \, dE$$

 $F^{1/2}$

et E/kT = x, then E=kT

Therefore dE = kTdx

$$I = 2\pi V \left(\frac{2m}{h^2}\right)^{\frac{3}{2}} \int_0^\infty (kTx)^{\frac{1}{2}} e^{-\frac{x}{kT}} (kT) dx$$

$$= 2\pi \mathrm{V} \left(\frac{2mkT}{h^2}\right)^2 \int_0^\infty (x)^{\frac{1}{2}} e^{-x} dx$$
$$= 2\pi \mathrm{V} \left(\frac{2mkT}{h^2}\right)^{\frac{3}{2}} \int_0^\infty (x)^{\frac{3}{2}-1} e^{-x} dx$$

The integral on the R.H.S of this equation is a gamma-function defined as

$$\int_0^\infty (x)^{n-1} e^{-x} dx = \Gamma (n)$$

Therefore $\int_0^\infty (x)^{\frac{3}{2}-1} e^{-x} dx = \Gamma(3/2) = \frac{1}{2} \sqrt{\pi}$

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I =
$$2\pi V \left(\frac{2mkT}{h^2}\right)^{\frac{3}{2}} x^{\frac{1}{2}} \sqrt{\pi} = V \left(\frac{2\pi mkT}{h^2}\right)^{\frac{3}{2}}$$

Sub. the value of this integral in eq. (2),

The constant A i.e., $e^{-\alpha}$ is called degeneracy parameter

Taking log on both sides,
$$-\alpha = \log \left[\frac{N}{V} \left(\frac{h^2}{2\pi m k T} \right)^{3/2} \right]$$
 ------(4)

Maxwell –Boltzmann speed Distribution Law

The Maxwell –Boltzmann equation for distribution of energy among the molecules of an ideal gas is

$$n(E)dE = \frac{2\pi N}{(\pi KT)^{3/2}} E^{1/2} e^{-E/KT} dE ------(1)$$

A classical ideal gas is defined as an assembly of non-interacting molecules, each distinguishable from the other. Therefore, the molecules have no internal degrees of freedom, all the energy of the gas in the form of kinetic Energy of the molecules.

$$E = \frac{1}{2}mv^{2} = \frac{1}{2}m(v^{2}_{x} + v^{2}_{y} + v^{2}_{z}) \qquad ------(2)$$

dE=mvdV ------(3)

Substitute the expression for E and dE in eq(1), the number of molecules n(v)dv whose speeds lie in between v and v + dv is given by

$$n(v)dv = \frac{2\pi N}{(\pi KT)^{3/2}} (\frac{1}{2}mv^2)^{1/2} e^{-mv^2/2KT} mv dv$$
$$= 4\pi N (\frac{m}{2\pi KT})^{3/2} v^2 e^{-mv^2/2KT} dv \qquad -----(4)$$

This eq. is known as the Maxwell of Maxwell –Boltsmann law of distribution speeds among the molecules of a gas . In this equation n(v) is the number of molecules per unit speed range .Therefore, the unit of n(v) is molecules /(m/sec).



Discussion of the law:

The curves for n(v) plotted against v at three different temperatures $T_1,\!T_2,\!T_3$ where $T_1\!\!<\!\!T_2\!\!<\!\!T_3$

From the distribution curves we get the following conclusions.

(1)At any temperature there is no molecules having zero speed.

(2) As the speed increases the no of molecules in a given speed interval Δv increases upto a certain maximum value.

(3) As the speed further increases beyond v_p , n(p) decreases exponentially towards zero. It means according to classical physical a molecules can have a infinite speed.

(4) As the temperature increases, v_p increases, and the range of speed is greater .Hence the curve become broad.

(5) At the given temperature the area under the distribution curves is equal to the total number of molecules in the gas .Thus

 $N=\int_0^\infty n(v)dv$

Since the area must be same at all the temperature, the distributive curve must flatten as the temperature rises.

Most Probable, Average and root mean square speed

Most Probable speed, v_p:

The most probable speed of the molecules is that speed at which the number of molecules per unit range of speed is maximum.

From the M-B distribution law for the molecular speeds the number of molecules per unit range of speed is given by

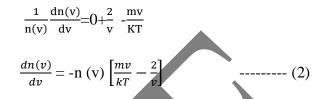
n (v) =
$$4\pi N \left(\frac{m}{2\pi kT}\right)^{3/2} v^2 e^{\frac{-mv^2}{2kT}}$$
 ------(1)



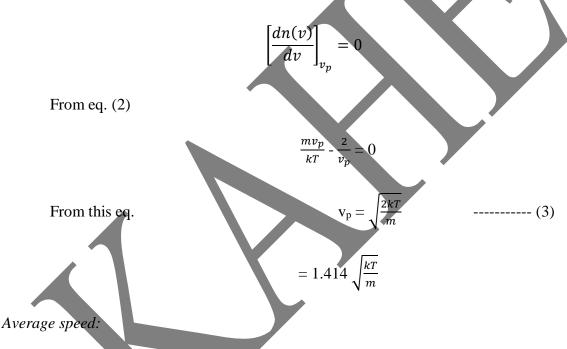
Taking logarithm of both sides of this equation

$$\log n (v) = \log \left[4\pi N \left(\frac{m}{2\pi kT} \right)^{3/2} \right] + 2 \log v - \frac{mv^2}{2kT}$$

by dif.this equation with the respect v,



At the most probable speed $v=v_p$ the number of molecules n(v) is maximum



The number of molecules whose speeds lie between v and v+dv is n (v) dv. The total speeds of these molecules is v n (v)dv, and the total number the molecules is N. Since the molecules is distributed among all velocities from 0 to ∞ , the average speed is given by

$$\tilde{v} = \frac{1}{N} \int_{0}^{\infty} v n(v) dv \qquad ----(4)$$
$$= \frac{1}{N} \int_{0}^{\infty} v 4\pi N \left(\frac{m}{2\pi kT}\right)^{3/2} v^{2} e^{\frac{-mv^{2}}{2kT}} dv$$
$$= 4\pi \left(\frac{m}{2\pi kT}\right)^{3/2} \int_{0}^{\infty} v^{3} e^{\frac{-mv^{2}}{2kT}} dv \qquad -----(5)$$

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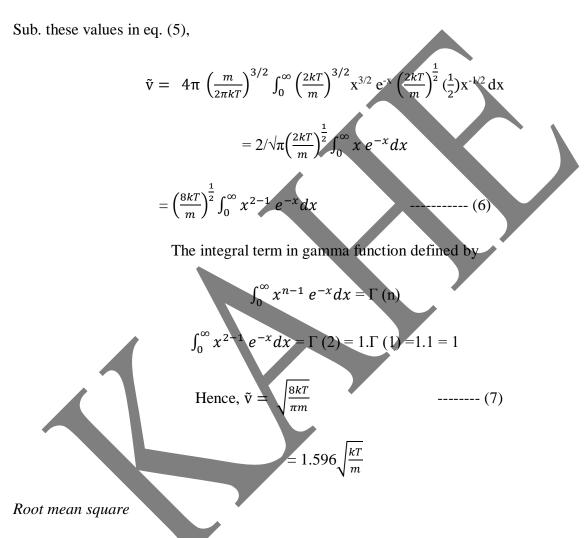


Let $mv^2 / 2kT = x$

Then

$$\mathbf{v} = \left(\frac{2kT}{m}\right)^{1/2} \mathbf{x}^{1/2}$$

$$\mathrm{d}\mathbf{v} = \left(\frac{2kT}{m}\right)^{\frac{1}{2}} \left(\frac{1}{2}\right) \mathbf{x}^{-1/2} \,\mathrm{d}\mathbf{x}$$



The number of molecules whose speed is between v and v + dv is n (v) dv. The sum of the squares of the speeds of these molecules is v^2 n (v) dv, and the total number of molecules is N. Since the total number of molecules is distributed among all these from 0 to ∞ , the mean square speed is given by

$$\tilde{v}^{2} = \frac{1}{N} \int_{0}^{\infty} v \, n \, (v) \, dv \qquad ------ (8)$$



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$$= \frac{1}{N} \int_0^\infty v^2 \ 4\pi N \ \left(\frac{m}{2\pi kT}\right)^{3/2} \ v^2 \ e^{\frac{-mv^2}{2kT}} dv$$

Let $mv^2 / 2kT = x$

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Then

$$\mathrm{dv} = \left(\frac{2kT}{m}\right)^{\frac{1}{2}} \left(\frac{1}{2}\right) x^{-1/2} \mathrm{d}x$$

Sub. these values in eq. (5),

 $\mathbf{v} = \left(\frac{2kT}{m}\right)^{1/2} \mathbf{x}^{1/2}$

The integral term in gamma function defined by

Thus $v_p < \tilde{v} < v_{rms}$

Principles of Equipartition of Energy:

The total energy of a particle of a system in thermodynamic equilibrium can be expressed as the sum of independent squared terms in position and momentum coordinates. For example,



the K.E. E of a free particle of mass , m which has velocity v and momentum p can be expressed as

$$E = \frac{p_x^2 + p_y^2 + p_z^2}{2m} \qquad ------(1)$$

For a particle of mass m moving with simple harmonic motion along the x-axis, the total energy

$$E = \frac{p_x^2}{2m} + \frac{1}{2}Cx^2 \qquad -----(2)$$

where C is the force constant per unit displacement from the mean position. Each independent squared term in the expression for the energy of a particle is said to give rise to one degrees of freedom of the particles. Thus a free particle has three degrees of freedom, and a particle moving with a linear simple harmonic motion has two degrees of freedom.

The principle of equipartition of energy is stated as follows:

When a system in a thermodynamic equilibrium at absolute temperature T, the mean value of each quadratic term in either a position or a momentum coordinate, which occurs in the total energy of the particle is (1/2)kT.

The principle may also be stated as follows:

When a system is in thermodynamic equilibrium at absolute temperature T, the mean energy of a particle in the system is distributed equally among its various degrees of freedom and for each of them it is (1/2)kT.

The principle was first deduced by Maxwell in 1959 for the energy of translational motion of a free particle. Boltzmann later showed that the principle is true for the energies of the rotation and vibration also. Rigorous proof's from statistical mechanics were given later by other workers.

Proof of the Principles:

We will prove the principle by finding the mean value of the term $\frac{p_x}{2m}$ in the expression for the energy of a particle in a linear S.H.M. along the X-axis.



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The mean value of $\frac{p^2}{2m}$ at equilibrium is given by

$$\left\langle \frac{p^2}{2m} \right\rangle = \frac{\int \int \int \int \int \frac{p^2}{2m} e^{-\frac{E}{kT}} dx dy dz dp_x dp_y dp_z}{\int \int \int \int \int e^{-\frac{E}{kT}} dx dy dz dp_x dp_y dp_z}$$

Where E is the total energy.

Writing the exponential term as

$$e^{-\frac{E}{kT}} = e^{-\frac{px^2}{2m} + Eo}/kT$$
$$= e^{-\frac{px^2}{2mkT} + e^{-Eo/kT}}$$

Where E_0 is the contribution to the total energy due to all the coordinates and momenta expect P_x, we get

$$\left\langle \frac{p_x^2}{2m} \right\rangle = \frac{\int \frac{p_x^2}{2m} e^{-\frac{p_x^2}{2mkT}} dp_x \int \int \int \int e^{-\frac{Eo}{kT}} dx dy dz dp_y dp_z}{\int e^{-\frac{p_x^2}{2mkT}} dp_x \int \int \int \int e^{-\frac{Eo}{kT}} dx dy dz dp_y dp_z}$$

Cancelling the five –fold integral in the numerator and denominator, we obtain

$$\left\langle \frac{p_{x}^{2}}{2m} \right\rangle = \frac{1}{2m} \frac{\int_{-\infty}^{\infty} p_{x}^{2} e^{-\frac{p_{x}^{2}}{2mkT}} dp_{x}}{\int_{-\infty}^{\infty} e^{-\frac{p_{x}^{2}}{2mkT}} dp_{x}} - \dots (3)$$
Let, $\frac{p_{x}^{2}}{2mkT} = u^{2}$
So that $p_{x} = \sqrt{(2mkT)} u$

$$dp_{x} = \sqrt{(2mkT)} du$$
hence, $\left\langle \frac{p_{x}^{2}}{2m} \right\rangle = \frac{1}{2m} \frac{\int_{-\infty}^{\infty} (2mkT) u^{2} e^{-u^{2}} \sqrt{(2mkT)} du}{\int_{-\infty}^{\infty} e^{-u^{2}} \sqrt{(2mkT)} du}$

$$= kT \cdot \frac{\int_{-\infty}^{\infty} u^2 e^{-u^2} du}{\int_{-\infty}^{\infty} e^{-u^2} du}$$
 ------ (4)

now evaluate the integral in the numerator.

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$$I = \int_{-\infty}^{\infty} u^2 e^{-u^2} \quad du = \int_{-\infty}^{\infty} u \cdot (u e^{-u})^2 \quad du$$

Integrating by parts, we obtain

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$$I = u \int u e^{-u^2} du - \int \left[\frac{du}{du} \int u e^{-u^2} du\right] du$$
$$= -\frac{1}{2} \left[ue^{-u^2}\right] + \frac{1}{2} \int_{-\infty}^{\infty} e^{-u^2} du$$

It can be shown that the first term is zero at both the limits. Hence $\int_{-\infty}^{\infty} u^2 e^{-u^2} du = \frac{1}{2}$

$$\int_{-\infty}^{\infty} e^{-u^2} du.$$
 Sub this in eq. 4, $\left\langle \frac{p_x^2}{2m} \right\rangle = \frac{1}{2} kT$

By a similar proof we can shoe that the mean value $(1/2)C x^2$ is (1/2) kT. Thus the mean energy of one- dimension harmonic oscillator is

$$\left<\frac{p_{x}^{2}}{2m}\right> + \left<\frac{1}{2}Cx^{2}\right> = \frac{1}{2}kT + \frac{1}{2}kT = kT$$

In the case of a free particle of the total K E

$$\frac{p_x^2 + p_y^2 + p_y^2}{2m}$$

The mean energy

$$\langle \frac{p_x^2}{2m} \rangle + \langle \frac{p_y^2}{2m} \rangle + \langle \frac{p_z^2}{2m} \rangle = \frac{1}{2}kT + \frac{1}{2}kT + \frac{1}{2}kT = \frac{3}{2}kT$$

Limitation of the principle of equipartition of Energy:

Theoretical values of the specific heat capacities of substances calculated from the equipartition of energy show that they should be independent of the temp. But the experimental result shows that the conclusion is not true. The effect of temperature on the specific heat capacity is considerable. The specific heat capacity increases with increase in temperature, and it decreases when the temperature is lowered. At low temperatures its rate of decrease with decrease of temperature is large both for solids and gases. The effect cannot be explained in any way by classical mechanics and the equipartition principle.



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Partition function Z

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Where $A = e^{-\alpha}$

The summation in eq. (1) is taken over all integrals of r corresponding to all possible energy states of the particles. The summation term is called the partition function or the sum over states and is denoted by the symbol Z. Thus g_re^{-E/k1} -- (2) Now in terms of Z, eq. (1) is written as $N = AZ = e^{-\alpha}Z$ (3)

Using this relation the Maxwell-Boltzmann distribution law $n_r = e^{-\alpha}g_r e^{-E/kT}$

Is written in the form

$$n_r = N/Z \times g_r e^{-Er/kT}$$
 ------(4)

----- (5b)

the multiplier α can be expressed as follows

$$e^{-\alpha} = N/Z$$

 $-\alpha = \log (N/Z) = -\log (Z/N)$ ----- (5a)

 $\alpha = -\log (N/Z) = \log (Z/N)$

Evaluation of Z

$$Z = N/A = \frac{N}{\frac{N}{V} \left(\frac{h^2}{2\pi m kT}\right)^{3/2}}$$

This equation gives the value of the partition function for an ideal gas consisting of mono atomic having no spin

Condition for applicability of the Maxwell-Boltzmann statistics

The M-B statistics is applicable to a system of identical particles which do not interact with each other directly except during their collisions. The condition whether neutral particles in a system will interact with each other or not determined by calculating their de Broglie wavelength.

If the wavelength is smaller than the mean distance between the particles, then the particles will not interact. Thus if λ is the de Broglie wavelength and d the mean distance between the particles, the condition for application of M-B statistics is $\lambda < d$ ------ (1)

To express this condition in terms of the degeneracy parameter $e^{-\alpha}$ the expression for d and λ .

The volume per particle = V/N

Therefore the mean distance d between the particles is

(2)

The mean K.E. of the particles is (3/2) kT. Therefore, the corresponding momentum p is given

 $d = (V/N)^{1/3}$

$$\frac{p^2}{2m} = \frac{3}{2}kT$$

$$p = \sqrt{(3mkT)}$$
and $\lambda = h/p = h / \sqrt{(3mkT)}$

$$= \left(\frac{h^2}{3mkT}\right)^{1/2} \qquad \dots \dots \dots (3)$$

Sub. the values of d and λ in condition (1),

by

$$\left(\frac{h^2}{3mkT}\right)^{1/2} < \left(\frac{V}{N}\right)^{1/3}$$

Or
$$\left(\frac{h^2}{3mkT}\right)^{3/2} < (V/N)$$



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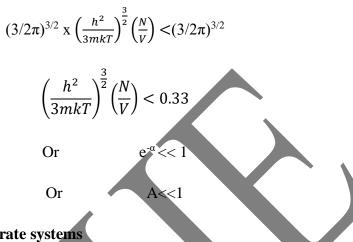
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$$\operatorname{Or}\left(\frac{h^2}{3mkT}\right)^{\frac{3}{2}}(N/V) < 1$$

Multiplying by $(3/2\pi)^{3/2}$,

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Non -degenerate and degenerate systems

If the number of particles, n_r in an energy level E_r is much less than the number of quantum states, g_r available in the same energy level i.e., if $n_r \ll g_r$ the system of particles is said to be non degenerate.

If n_r is greater than g_r i.e., if $n_r > g_r$ the system is said to be degenerate.

If n_r is much greater than g_r , i.e., if $n_r >> g_r$ the system is said to be strongly degenerate.

The degeneracy parameter $A(e^{-\alpha})$ is given by

If we consider the ground state as a zero, then



Therefore in terms of A is foregoing conditions are as follows:

 $A=e^{-\alpha}=$

1. If A<<1, the system is non-degenerate.

2. If A>1, the system is degenerate.

3. If A >>1, the system is strongly degenerate.

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Maxwell velocity distribution in a given direction

In a system of an ideal gas the number of molecules having velocity components in the range between v_x and v_x+dv_x , v_y and v_y+dv_y and v_z and v_z+dv_z is given by

$$n(v_x, v_y, v_z) dv_x dv_y dv_z = [f(v_x, v_y, v_z)] [g(v_x, v_y, v_z) dv_x dv_y dv_z]$$
 ------(1)

The first term on the R.H.S of this equation is the distributed function for the velocity components, i.e., it is the number of molecules each having the velocity components $v_x v_y v_z$ per quantum state in the energy level E. The second term is the number of quantum states within the velocity space $dv_x dv_y dv_z$. These terms are obtained as follows.

f (E) =
$$\frac{N}{V} (\frac{h^2}{2\pi m KT})^{3/2} e^{-E/kT}$$

Assuming that each molecule has only three degrees of freedom due to its motion of translation, the K.E of each molecule is given by

 $E = \frac{1}{2} mv^2$

From eq. (2),
$$f(v_x, v_y, v_z) = \frac{N}{V} \left(\frac{h^2}{2\pi m k T}\right)^{3/2} e^{-m(v_x 2 + v_y 2 + v_z 2)/kT}$$
 -----(4)

The volume of one quantum state in the momentum space is h^3/V , where V is the physical volume of the system.

Therefore the number of quantum states in volume $dp_x dp_y dp_z$ of momentum space.

$$= V/h^3 dp_x dp_y dp_z$$
$$= m^3 V/h^3 dv_x dv_y dv_z$$

Hence g (v_x, v_y, v_z) dv_x dv_y dv_z =
$$m^3 V/h^3 dv_x dv_y dv_z$$
 ------ (5)

Sub. Eq. (4) and (5) into (1)

n (v_x, v_y, v_z) dv_x dv_y dv_z = $\frac{N}{V} \left(\frac{h^2}{2\pi m KT}\right)^{3/2} m^3 V/h^3 e^{-m (v_x^2 + v_y^2 + v_z^2)/kT} dv_x dv_y dv_z$



Simplifying this equation,

n (v_x, v_y, v_z) dv_x dv_y dv_z =
$$N \left(\frac{m}{2\pi KT}\right)^{3/2} x e^{-m (v_x^2 + v_y^2 + v_z^2)/kT} dv_x dv_y dv_z$$
 ------(6)

Now the number n (v_x) dv_x of molecules, having x component of velocity in the range between v_x and v_x +dv_x is obtained by integrating eq. (6) over all possible values of v_y and v_z .

n (v_x) dv_x =
$$\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} [n (v_x, v_y, v_z) dv_x] dv_y dv_z$$

= $N \left(\frac{m}{2\pi KT}\right)^{3/2} e^{-mv_x^2/2kT} \int_{-\infty}^{\infty} e^{-mvy^2/2kT} \int_{-\infty}^{\infty} e^{-mvy^2/2kT}$

The definite integrals are standard integrals, the value of each being

 $[(2\pi kT)/m]^{1/2}$

Sub. this value in above equation,

n (v_x) dv_x =
$$N \left(\frac{m}{2\pi KT}\right)^{3/2} \left(\frac{2\pi KT}{m}\right) e^{-mv_x^2/2kT} dv_x$$

= $N \left(\frac{m}{2\pi KT}\right)^{1/2} e^{-mv_x^2/2kT} dv_x$ (7)

This equation gives the number of molecules having x component of velocity in the range between v_x and v_x +dv_x. Eq. (7) shows that the velocity component v_x is distributed symmetrically about the value $v_x = 0$.

Total internal energy of an ideal gas

The total internal energy of an ideal gas is given by

$$\mathbf{U} = \int_{0}^{\infty} E \ n(E) dE \qquad ------ (1)$$

$$= \int_{0}^{\infty} E \left[\frac{2\pi N}{(\pi kT)^{3/2}} E^{\frac{1}{2}} e^{-\frac{E}{kT}} dE \right]$$

$$= \frac{2\pi N}{(\pi kT)^{3/2}} \int_{0}^{\infty} \left[E^{\frac{3}{2}} e^{-\frac{E}{kT}} dE \right] \qquad ------ (2)$$

Let E/kT = x, so that E = kTx



Therefore dE = kTdx

Sub. These values in eq. (2) get

$$= \frac{2\pi N}{(\pi kT)^{3/2}} (kT)^{3/2} kT \int_0^\infty x^{3/2} e^{-x} dx$$

$$= \frac{2NkT}{\sqrt{\pi}} \int_0^\infty x^{\frac{5}{2}-1} e^{-x} dx \qquad -----(3)$$

The integral is the gamma function

$$\Gamma$$
 (5/2) = 3/2 Γ (3/2) = 3/2 x $\sqrt{\pi}$ = 3 $\sqrt{\pi}$ / 4

Sub. This value in eq. (3),

$$U = \frac{2NkT}{\sqrt{\pi}} \times \frac{3\sqrt{\pi}}{4} = \frac{3}{2} NkT$$

From this equation, the average internal energy per molecule is given by U/N = $\frac{3}{2}$ kT ----- (5)

For 1 mole of an ideal gas, N is Avagadro's number and the value of U is given by eq. (4) is the total internal energy of one mole of an ideal gas.

Molar Heat Capacity of a gas at Constant Volume

The molar heat capacity C_v , of a gas at constant volume is defined as the quantity of heat required to raise the temperature of 1 mole of the gas through 1 degree, at constant volume.

According to the definition

$$C_{v} = \left(\frac{\partial U}{\partial T}\right) v$$
$$= \frac{\partial}{\partial T} \left(\frac{3}{2}NKT\right) = \frac{3}{2}NK$$
$$C_{v} = \frac{3}{2}N\left(\frac{N}{R}\right) = \frac{3}{2}R$$

-----(1)

$$N(\frac{N}{R}) = \frac{3}{2}R \qquad -----(2)$$

Where R is the gas constant for one mole.



Entropy

According to Boltzmann's relation, the entropy S of an isolated system of non-interacting particles in equilibrium is given by

Where W_{max} is the maximum number of statistically independent ways of distributing the particles among the quantum states. From the M-B count,

$$\begin{split} \log W &= N \log N - N + \Sigma_r n_r \left(\log g_r - \log n_r + 1 \right) \\ &= N \log N - N + \Sigma_r n_r \log \left(g_r / n_r \right) + N \\ &= N \log N - \Sigma_r n_r \log \left(n_r / g_r \right) \end{split}$$

For maximum value of W

$$n_r/g_r = e^{-(\alpha+1)}$$

Therefore, $\log W_{max} = N \log N + \Sigma_r n_r (\alpha + \beta E_r)$

= N log N + $\alpha \Sigma_r n_r + \beta \Sigma_r n_r E_r$

 $= N \log N + \alpha N + \beta U \qquad -----(2)$

Where U is the total internal energy

Sub. $\beta = 1/kT$ and $\alpha = \log (Z/N)$

Where Z is the partition function,

 $\log W_{max} = N \log N + N \log \frac{Z}{N} + \frac{U}{kT}$

Sub. this value of log W_{max} in eq. (1),

$$S = U/T + Nk \log Z \qquad -----(4)$$

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For an ideal gas, $U = \frac{3}{2} \text{ NkT}$ and $Z = V \left(\frac{2\pi m KT}{h^2}\right)^{\frac{3}{2}}$

According to M-B count, the entropy of an ideal gas is

Helmholtz Free Energy

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COURSE CODE: 19PHP201

The Helmholtz Free Energy F of a system of particles is defined by

F=U-TS

The entropy of a system is given by

$$S = \frac{U}{T} + NK + NK \log \frac{Z}{N}$$

Subs this value in eq (1)

$$F=U-T[\frac{v}{r}+NK+NK\log\frac{z}{N}]$$

=-NKT $(log \frac{Z}{N}+1)$

=-NKT[log $Z - \log N + 1$] ------ (3)

(2)

Eq(3) can be expressed in

 $F = -NKT \log Z + NKT \log N - NKT$

-NKT $\log Z$ +KT(N $\log N - N$)

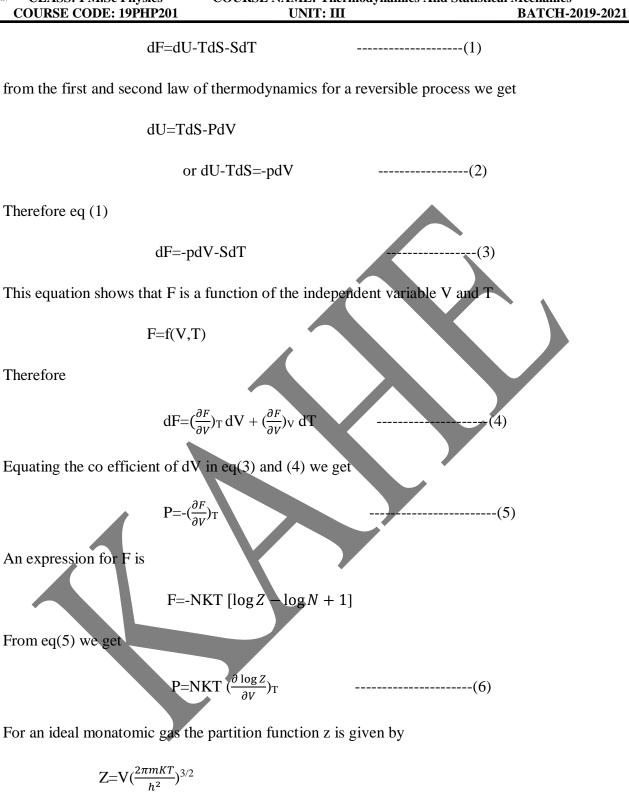
=-NKT $\log Z$ +KT $\log N$!

Pressure and Equation of State of an Ideal gas

The Helmholtz free energy F is

F=U-TS

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By taking log to base e

$$\log Z = \log V + \frac{3}{2} \log(\frac{2\pi m KT}{h^2})$$

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By diff. partially we get

$$\frac{\partial \log Z}{\partial V} = \frac{1}{V}$$

subs this value in eq(6)

$$P = \frac{NKT}{V}$$

-----(7)

(8)

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In this equation K is Boltzmann's constant (1.38x10⁻²³J/K), and for 1 mole of a gas N is Avogardro's number (6.023 $\times 10^{23}$ /mol). Therefore, for one mole the product NK is the same for all gases. This constant is called molar gas constant denoted by the symbol R. So for 1 mole of a gas eq(7) in the form

PV=RT

This equation for one mole of an ideal gas is called the equation of state for the gas. The numerical value of R is

R=Nk=(6.023 X10²³/mol)(1.38 x10⁻²³ J/K)

=8.31 J/mol K

Limitations of Maxwell-Boltzmann Method

The method has several limitations, some of them are

- 1. It is applicable to only isolated gas of identical molecules in equilibrium, for which the following conditions are satisfied.
 - i. The mean potential energy due to interaction b/w the molecules are very small compared to their mean K E.
 - The gas is dilute i.e, the number of molecules per unit volume is small, so the ii. average separation between the molecules is large and hence individual molecules can be distinguished.

Therefore, important results, such as the expression for u, c_v and p, obtained by this method are the same as those delivered by applying a simple kinetic theory.



2. The expression for the Maxwell-Boltzmann count does not give the correct expression for the entropy of an ideal gas, and leads to Gibbs'paradox. To resolve the Paradox the expression must be divided by N!.

UNIT: III

3. In the expression for the entropy of an ideal gas

$$S=NK\log[(\frac{2\pi mKT}{h^2})e^{5/2}]$$

if we put T=0, we get

S=NKlog 0=- ∞ .

Thus the expression for S does not satisfy the third law of thermodynamics which may be started as follows: Every substance has a finite positive entropy, but an absolute zero of temperature the entropy may become zero, and it becomes zero in case of perfectly crystalline solid.

 It cannot be applied to a system of indistinguishable particle. If we apply the Maxwell – Boltzmann distribution law of thermionic emission, we get the following expression for the emission current density

 $J = A_0 T^{1/2} e^{-\frac{3}{K}}$

which is not correct .The correct Expression which has been verified as

 $J=A_cT^2e^{-\frac{\varphi}{\kappa T}}$

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Possible Questions

- 1. What is meant by partition function?
- 2. What are called boson?
- 3. Define molar heat capacity of a gas.
- 4. Define degenerate and non-degenerate system.
- 5. Explain the principle of equipartition of energy.
- 6. Derive an expression for pressure and the equation of state of an ideal gas.
- 7. What are the limitations of M-B statistics?
- 8. Obtain Helmhotz free energy.
- 9. Obtain an expression for most probable, average and root mean square speeds.
- 10. Explain the limitations of M.B. method.
- 11. Calculate total internal energy of an ideal gas.
- 12. Evaluate constant value of M.B. statistics.
- 13. Explain M.B. distribution law.
- 14. What are the conditions for applicability of M.B. statistics?

Coimbatore-641021.

(For the candidates admitted from 2018 onwards)

DEPARTMENT OF PHYSICS

UNIT III (Objective Type/Multiple choice Questions each Questions carry one Mark)

THERMODYNAMICS AND STATISTICAL MECHANICS

TART -A(Onine Examination)							
S.No.	QUESTIONS	OPTION 1	OPTION 2	OPTION 3	OPTION 4	ANSWER	
1	Which is called as degeneracy parameter?	e ^α	e ^β	ce ^{-a}	$e^{+\beta}$	e ^{-α}	
2	Partition function is denoted by the symbol	Z	А	М	N	Z	
3	Mean distance d between the particles is	$(N/V)^{1/3}$	$(V/N)^{1/3}$	$(N/V)^3$	$(V/N)^3$	$(V/N)^{1/3}$	
4	If A<<1, the system is	non- degenerate	degenerate	strongly degenerate	weekly degenerate	non- degenerate	
5	If A>1, the system is	non- degenerate	degenerate	strongly degenerate	weekly degenerate	degenerate	
6	If A>>1, the system is	non- degenerate	degenerate	strongly degenerate	weekly degenerate	strongly degenerate	
7	The number of quantum states is represented by	n _r	gr	Er	N	gr	
8	The degeneracy parameter A =	Z	e ^α	e ^{-α}	e^{β}	e ^{-α}	
9	In M.B. distribution, the unit of n(v) is	m/ sec	mol/sec	mol/m/sec	sec	mol/m/sec	
10	v _p =	$\sqrt{(2kT/m)}$	1.414√ kT	1.732√(kT/m)	3.414√kT	$\sqrt{(2kT/m)}$	
11	Average speed is represented by	Vp	v	Av	V _{rms}	Vp	
12	Which of the following is correct for a perfect gas?	v <vpvrms< td=""><td>v_p<v <v<sub="">rms</v></td><td>$v < v_{rms} < v_p$</td><td>v>v_p>v_{rms}</td><td>v_p<v <v<sub="">rms</v></td></vpvrms<>	v _p <v <v<sub="">rms</v>	$v < v_{rms} < v_p$	v>v _p >v _{rms}	v _p <v <v<sub="">rms</v>	

PART –A(Online Examination)



13	Root mean square =	$\sqrt{(2kT/m)}$	1.414√ kT	1.732√(kT/m)	3.414√kT	1.732√(kT/m)
14	Average speed =	1.596√(kT/m)	1.414√ kT	1.732√(kT/m)	3.414√kT	1.596√(kT/m)
15	The equation for total internal energy of one mole of an ideal gas is	$U/N = 3/2 \ kT$	U/N = 3/2	U/N = kT	$U/N = 2/3 \ kT$	U/N = 3/2 kT
16	The value of $\beta =$	kT	k	Т	1/kT	1/kT
17	At absolute zero temperature the entropy may become The value of entropy becomes zero in perfectly	Infinity	positive	Zero	Negative	Zero
18	The value of entropy becomes zero in perfectly	liquid	crystalline solid	gas	inert gas	crystalline solid
19	M.B. law cannot be applicable to particles.	distinguishable	Indistinguishable	isolated	Isobaric	Indistinguishable
20	In dilute gas, the number of molecules per unit volume is	large	very small	infinity	small	small
21	In dilute gas, the average separation between the molecules is	large	very small	Infinity	small	large
22	The mean energy of principle of equipartition of	kT	3/2 kT	2/3 kT	¹∕2 kT	3/2 kT
23	A free particle has degrees of freedom.	1	2	3	4	3
24	A particle moving with a linear simple harmonic motion has degrees of freedom.	1	2	3	4	2
25	Helmhotz free energy F of a system of particles is defined by	F=U-TS	F=U/TS	F=U+TS	F=UTS	F=U-TS
26	The specific heat at constant volume is	the amount of heat required to raise the temperature of unit mass of gas through one degree, at constant pressure	the amount of heat required to raise the temperature of 1 kg of water through one degree	the amount of heat required to raise the temperature of unit mass of gas through one degree, at constant volume	any one of the above	the amount of heat required to raise the temperature of unit mass of gas through one degree, at constant volume
27	The gas constant (R) is equal to the of	sum	difference	product	ratio	difference

	two specific heats.					
28	The quantum statistics reduces to classical statistics under the following condition	$\rho\lambda^3 = 1$	$\rho\lambda^3 \gg 1$	$\rho\lambda^3 \ll 1$	$\rho\lambda^3 = 0$	$\rho\lambda^3 >> 1$
29	Specific heat of metals can be expressed as	T^3	$AT + BT^2$	AT^2+BT^3	$AT + BT^3$	$AT + BT^3$
30	Boltzmann entropy probability relation is given by	S=k log _e ω	$S = k/log_e\omega$	$S = k {+} log_e \omega$	$S = k - log_e \omega$	S=k log _e ω
31	Enthalpy and internal energy have relation	H = U - PV	H = U/PV	H=U+PV	H=UPV	H=U+PV
32	In quantum physics identical particles are	a) indistinguishable	distinguishable	symmetric	anti-symmetric	indistinguishable
33	The zero point energy of one dimensional oscillator is	2h	½ h	1/3 h	3h	¹⁄₂ h
34	In classical physics identical particles are	indistinguishable	distinguishable	symmetric	anti-symmetric	distinguishable
35	The dimensions of the phase space depends upon the of the system.	entropy	heat content	degrees of freedom	enthalpy	degrees of freedom
36	of a system of particles is given by $F = U - TS$	Helmholtz free energy	free energy	helmholtz function	Gibb's free energy	Helmholtz free energy
37	For non-degenerate system	A= 1	A<< 1	A> 1	A>= 1	A<< 1
38	The spin of the photon is	0	1	2	1/2	1
39	B.E distribution function is given by	$\{1/(e^{a+bE})\}$	$\{1/(e^{a+bE})+1\}$	$\{(e^{a+bE})-1\}$	{1/(ea + bE) - 1}	$\{(e^{a+bE})-1\}$
40	The degeneracy parameter $e^{-a} =$	$N/V (h^2/2pmkT)^{1/2}$	N/V (h^2 / 2pmkT) ^{3/4}	$N/V (h^2 / 2pmkT)^{3/2}$	$N/V (h^2/2pmkT)^3$	$N/V (h^2 / 2pmkT)^{3/2}$
41	Maxwell first developedtheory	Equipartition	partition	classical	quantum	classical
42	According to classical mechanics a molecule can have	finite speed	infinite speed	variable speed	constant speed	infinite speed
43	As temperature increases, the most probablealso increases	frequency	wavelength	energy	velocity	velocity
44	B.E distribution law is used to deriveof radiation	Plank's law	Weiss law	Widemann- Franz law	All the above	Plank's law
45	Wave function of the system of identical Bosons is	Asymmetric	linear	non-linear	symmetric	symmetric
46	M.B. distribution can be applicable to	identical	indistinguishable	gas	liquid	identical

		molecule	molecule			molecule
47	In M.B. distribution the mean P.E. is than/ to K.E. of ideal gas.	larger	very large	small	equal	small
48	When T=0, the value of entropy $S = _$ in M.B. distribution.	infinity	negative infinity	zero	one	negative infinity
49	The correct expression for J =	$A_C T^2 e^{-\phi/kT}$	$A_C Te^{-\phi/kT}$	$A_C T^{1/2} e^{-\phi/kT}$	$T^2 e^{-\phi/kT}$	$A_C T^2 e^{-\phi/kT}$
50	The value of gas constant R=	8.13K/mol	7.013 mol/K	8.31 mol/JK	8.31 J/mol K	8.31 J/mol K
51	Partition function is denoted by the symbol	Ζ	А	М	Ν	Ζ
52	Mean distance d between the particles is	$(N/V)^{1/3}$	$(V/N)^{1/3}$	$(N/V)^3$	$(V/N)^3$	$(V/N)^{1/3}$
53	In M.B. distribution, the unit of n(v) is	m/ sec	mol/sec	mol/m/sec	sec	mol/m/sec
54	v _p =	$\sqrt{(2kT/m)}$	1.414√ kT	1.732√(kT/m)	3.414√kT	$\sqrt{(2kT/m)}$
55	Root mean square =	$\sqrt{(2kT/m)}$	1.414√ kT	1.732√(kT/m)	3.414√kT	1.732√(kT/m)
56	The equation for total internal energy of one mole of an ideal gas is	U/N = 3/2 kT	U/N = 3/2	U/N = kT	$U/N = 2/3 \ kT$	$U/N = 3/2 \ kT$
57	M.B. law cannot be applicable to particles.	distinguishable	Indistinguishable	isolated	Isobaric	Indistinguishable
58	The mean energy of principle of equipartition of energy is	kT	3/2 kT	2/3 kT	¹∕2 kT	3/2 kT



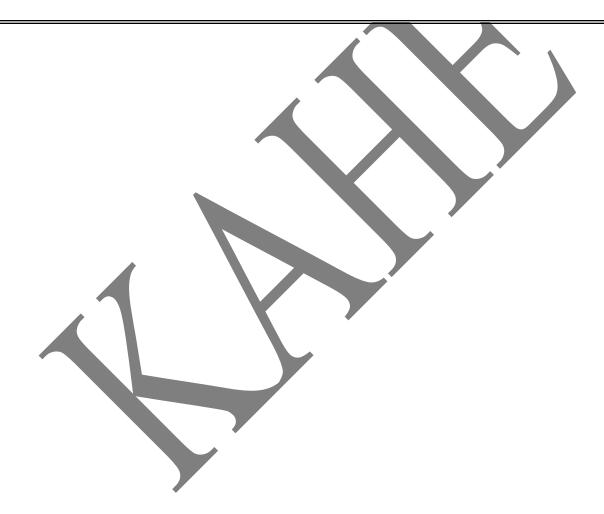
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UNIT-IV

Quantum Statistical Mechanics: B.E energy distribution for energies in the range E to E + dE - C ondition for B.E distribution to approach classical M.B distribution - Bose temperature - Bose Einstein condensation - Planck's law from B.E law - Fermi Dirac distribution law (no derivation) - FD law for the energies in the range E to E+dE - Fermi energy - Effect of temperature - Energy distribution curve - Free electron in a metal - Fermi temperature and Thermionic emission - Richardson Dushmann Equation - Comparison of MB,BE and FD statistics.



UNIT: IV

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Bose Einstein energy distribution for energies in the range E to E + dE

The molecules of an ordinary gas have spin angular momentum equal to an integral multiple of ħ. It means that the molecules are bosons and they will obey the Bose-Einstein statistics. The energy distribution law for a system of identical molecules is obtained as follows.

The number n (E) dE of the molecules having energies in the range from E to E+dE is given by

$$n(E) dE = f(E) g(E) dE$$
 ----- (1)

where f(E) is the energy distribution function, and g (E)dE is the number of quantum states available in the energy range.

Substituting the expression for f (E) in equation (1), obtained

$$n(E)dE = \frac{g(E)dE}{e^{\alpha}e^{E/kT}}$$

g (E) dE is given by, g (E) dE =
$$2\pi V (\frac{2m}{h^2})^{3/2} E^{1/2} dE$$

Substituting this equ. in (2)

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n(E) dE =
$$2\pi V \left(\frac{2m}{h^2}\right)^{3/2} \cdot \frac{E^{\frac{1}{2}}dE}{e^{\alpha}e^{E/kT}-1}$$
 ------ (3)

This is Bose Einstein energy distribution law given by the number of particles with energies between E and E + dE.

The constant e^{α} appears in the distribution law cannot be less than or equal to 1 because

(i) if e^{α} is <1, then for E = 0

$$n(E)dE = \frac{g(E)dE}{e^{\alpha}e^{E/kT} - 1}$$

$$n(0)dE = \frac{g(0)dE}{e^{\alpha}-1} = \text{negative quantity}$$

which is impossible.

(ii) if $e^{\alpha} = 1$, then for E = 0

$$n(0)dE = \frac{g(0)dE}{1-1} = \infty$$

Prepared by Dr.S.Sharmila, Assoc. Prof., Department of Physics, KAHE

Which is also impossible. Therefore e^{α} must be greater than 1.

Condition for B-E distribution to approach classical M-B distribution:

The B-E distribution is given by

n(E) dE =
$$2\pi V \left(\frac{2m}{h^2}\right)^{3/2} \cdot \frac{E^{\frac{1}{2}}dE}{\frac{1}{A}e^{E/kT} - 1}$$
 ------(1)

If 1 in the denominator is neglected in comparison with the first term, this distribution will approach the M-B distribution:

n(E) dE =
$$2\pi V \left(\frac{2m}{h^2}\right)^{3/2} AE^{1/2} e^{-E/kT} dE$$

= $2\pi V \left(\frac{2m}{h^2}\right)^{3/2} \frac{N}{V} \left(\frac{h^2}{2\pi m kT}\right)^{3/2} E^{1/2} e^{-E/kT} dE$
= $\frac{2\pi N}{(\pi kT)^3/2} E^{1/2} e^{-E/kT} dE$
The condition for this is that
 $\frac{1}{A} e^{E/kT} >> 1$
i.e. $\frac{A}{eE/kT} e^{E/kT} << 1$
Ill values of the energy, $e^{E/kT}$ is greater than or equal to 1. Therefore the condition

ndition is For

A<<1.

Limiting Case of Bose-Einstein Statistics

For an ideal Bose-Einstein distribution the degeneracy parameter $A(=e^{\alpha})$ cannot be greater than 1; its maximum value can be 1. If the temperature of the gas is decreased, the value of A increases from a low value towards 1.At a certain temperature T_{B} , the value of A becomes just less than 1, and then there is no change in the value below T_B . At T_B some proportion of the molecules start reaching the zero-energy state. This critical temperature is called the Bose temperature.



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Expression for the Bose Temperature:

In the ideal Bose-Einstein gas of spin less molecules in the thermal equilibrium at the temperature T, the tot al number of N molecules is given by

$$N = 2\pi V (\frac{2mkT}{h^2})^{3/2} \int_{0}^{2} \frac{x^{1/2}}{e^{\alpha} e^{x} - 1} dx$$
Where
$$x = E/kT$$
The value of integral
$$N = 2\pi V (\frac{2mkT}{h^2})^{3/2} \int_{2}^{\sqrt{\pi}} \left[\frac{1}{e^{\alpha}} + \frac{1}{\frac{3}{2}e^{2\alpha}} + \frac{1}{\frac{3}{2}e^{3\alpha}} + \dots \right]$$
Substituting the value $e^{1/A} = e^{-\alpha} = A$

$$N = V (\frac{2\pi mkT}{h^2})^{3/2} \left[\frac{A}{1} + \frac{A^2}{2^2} + \frac{A^3}{3^2} + \dots \right]$$
At Bose Temperature T=TB,
$$N = V (\frac{2\pi mkTB}{h^2})^{3/2} \left[1 + \frac{1}{2^2} + \frac{1}{3^2} + \dots \right]$$

The series in the square bracket is the Riemann Zeta function whose value is 2.612.

N=V
$$\left(\frac{2\pi m k T B}{h^2}\right)^{3/2}$$
 x 2.612
=2.612 V $\left(\frac{2\pi m k T B}{h^2}\right)^{3/2}$ -----(1)

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Where

For all known Bose-Einstein gas, the Bose temperature T_B is very low. For ex, for helium $(_2\text{He}^4)T_B=3.15~\text{K}$

Bose-Einstein condensation:

At the Bose temperature molecules T_B , molecules just start reaching the zero-energy state (E=0) from the higher energy state (E>0). If the temperature of the gas is lowered below T_B , the number of molecules in the zero-energy state will increase, and the number in the higher energy states will decrease. Suppose that in thermal equilibrium at the temperature T< T_B , n_0 is the number of molecules in the non zero- energy state and N_e is the number in the higher energy states. Then

where T<T_B

From eq.(3) & eq.(6) $n_0 = N - N_e$



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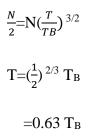
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 $=N[1-(\frac{T}{TB})^{3/2}]$

This eq. shows that as T approaches absolute zero of temperature $N_e \rightarrow 0$ and $n_0 \rightarrow N$

Thus all the molecules of B-E gas tend to condense into the zero energy states of the gas at absolute zero this phenomenon is called Bose- Einstein condensation. The temperature at which $n_0 = N_e = N/2$ obtained by sub. $N_{e=} N/2$ in eq. (6).



The molecules of an ideal B-E gas exist in two phases at $T < T_B$

- a gaseous phase consisting of Ne molecules distributed among the energy states higher i. than the ground state, and
- a condensed phase consisting of n_0 molecules occupying the ground state. ii.

The molecules in the condensed phase do not contribute to the internal energy, specific heat capacity, entropy, etc.

The transition of the molecules at T_B to the ground state is a sudden of phenomenon. In this phenomenon there is a decrease in volume of the momentum space by the volume of the space which n₀ molecules had occupied before their transition to the zero-energy state. This differs from the usual type of vapor condensation process in which there is a decrease in the physical volume.

The transition of liquid ₂He⁴-I to superfluid liquid ₂He⁴-II at the observed temperature 2.18K can be expanded by Bose-Einstein condensation process.



Planck's law of radiation from Bose-Einstein distribution law

According to planck's radiation law the energy of radiation of wavelength in the range between λ and λ +d λ emitted per unit volume by a perfectly black body at absolute temperature T

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is given by U (
$$\lambda$$
)d $\lambda = \frac{8\pi hc}{\lambda^5} \frac{d\lambda}{e^{\frac{hc}{kT}-1}}$

Derivation of law

According to the Bose- Einstein distribution law the number of Bosons having energies between E and E+dE is

$$n(E)dE = \frac{g(E)dE}{e^{\alpha}e^{\frac{E}{kT}}} \quad \dots \dots \quad (1)$$

where g(E) dE is the number of quantum stage of energies E and E +dE.

Let T be the absolute temperature of a black body chamber of volume v. The chamber is supposed to be filled with photons each having energy $h\nu$. They move in all possible directions with the speed of light C. Each photon has unit spin angular momentum equal to $h(h/2\pi)$. Hence photons are bosons and use B.E distribution law to derive Planck's law of radiation.

1. Constant α :

Photons of different energies are absorbed and re-emitted by the walls of the chamber at constant temp.

In this process a higher energy photon is converted into a number of low energy photons and vice-versa. Though the total energy of the photons remain constant, the total no of photons present in enclosure is not constant. Therefore the condition $\sum_r n_r = N$ or $\sum_r dn_r = 0$ is not applicable for the distribution and hence the multiplier α is zero, i.e e^{- α} is equal to 1.

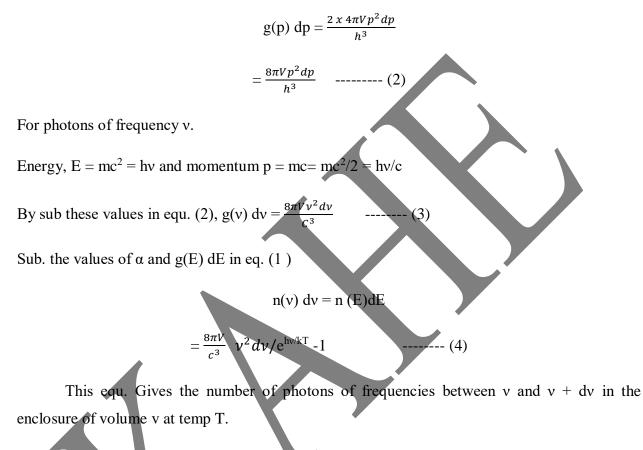
2. Expression for g(E)dE:

The number of quantum states corresponding to the momenta in the range between p and p+dp for particles with no spin is given by

$$g(p) dp = \frac{4\pi V p^2 dp}{h^3}$$

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each photon has unit spin angular momentum, there are two allowed quantum states for each photon. Hence photons of the same energy can have the two different directions of polarization circularly clockwise and circularly anti-clockwise. Taking the spin in to account for photons we have



Now the energy per unit volume of the enclosure, of the photons of frequencies between v and v + dv is

$$U(v)dv = \frac{hvn(v)dv}{v}$$
$$== \frac{8\pi h}{c^3} v^3 dv/e^{hv/kT} -1 \qquad -----(5)$$

This is this the plancks law of radiation in terms of frequency $\boldsymbol{\nu}$

Eqn(5) can be transformed in terms of the wavelength λ by using the relations

$$v = \frac{c}{\lambda}, dv = -\frac{c}{\lambda^2} d\lambda$$

The energy u(v)dv contained in a frequency interval between v and v + dv is equal to the contained in a corresponding wavelength interval $b/n \lambda$ and $\lambda + d \lambda$.

U(λ)d λ =u(ν)d ν

$$=\frac{8\pi h}{c^3}\frac{\left(\frac{c}{\lambda}\right)3\left(-\frac{c}{\lambda^2}d\lambda\right)}{eh\nu/kT-1}$$

Omitting the negative sign we get,

$$U(\lambda)d \lambda = \frac{8\pi hc}{\lambda^5} d\lambda / e^{h\nu/kT} - 1 - \cdots$$

This is the planck's law of radiation in terms of wavelength λ

Fermi –Dirac distribution law

In F-D statistics, the condition are:

 The particles are indistinguishable from each other i.e., there is no restriction between different ways in which n_i particles are chosen.

(6)

- (ii) Each sublevel or cell may contain 0 or one particle. Obviously g_i must be greater than or equal to n_i .
- (iii) The sum of energies of all particles in the different quantum groups taken constitutes the total energy of the system.

The Fermi – Dirac statistics is given by

$$n_i = \frac{g_i}{e^{(\alpha + \beta \epsilon i)} + 1}$$

Fermi-Dirac energy distribution for energies in the range E to E + dE

The number of particles having energies in the range between E and E+dE is given by

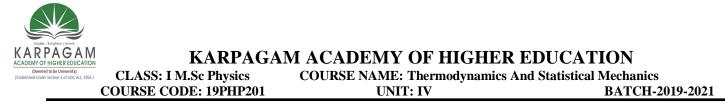
$$n(E) dE = f(E) g(E) dE$$
 ------(1)

where g(E) dE is the number of quantum states of energy between E and E +dE.

Sub. The expression for f(E) in eq. (1), then

$$n(E)dE = \frac{g(E)dE}{e^{\frac{Er-Ef}{kT}+1}} \qquad -----(2)$$

Prepared by Dr.S.Sharmila, Assoc. Prof., Department of Physics, KAHE



For particles like electrons of spin angular momentum $\pm 1/2$ ħ, there are two possible spin orientation. For a system of such particles g(E) dE is given by

g (E) dE = 2 x
$$2\pi V \left(\frac{2m}{h^2}\right)^{3/2} E^{1/2} dE$$

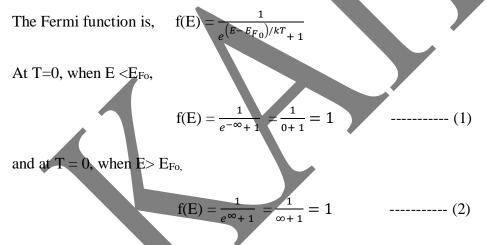
sub. the value of g (E) dE in eq. (2), get

n (E) dE =
$$4\pi V \left(\frac{2m}{h^2}\right)^{3/2} \frac{E^{\frac{1}{2}} dE}{e^{(E-E_F)/kT} + 1}$$
 (3)

This is Fermi-Dirac distribution law giving the number of particles with energies between E and E+dE.

Fermi Energy

The Fermi energy at absolute zero of temperature is denoted by E_{Fo} and this is considered as a constant over a large range of temperature.



Equ. 1 and 2 shows that at T=0, the function f (E) is constant equal to 1 for all values of energies upto E_{Fo} it falls to zero. That is at T=0, it is a step function.

Thus at absolute zero of temperature all possible quantum states of energy less than E_{Fo} re occupied and all those of energy more than E_{Fo} are empty.

Accordingly the Fermi energy E_{Fo} is defined as the energy of the highest occupied level at absolute zero. At any other temperature T>0, when $E = E_{Fo}$, the Fermi energy is



UNIT: IV

$$f(E) = \frac{1}{e^{\left(E - E_{F_0}\right)/kT} + 1} = \frac{1}{e^{0/kT} + 1} = 1/2$$

This means that at temperature T>0, the probability for occupation of a quantum states at the Fermi level is $\frac{1}{2}$. At temperature T>0, 50% of the quantum states at the Fermi level are occupied and 50% are empty.

Effect of temperature on Fermi Energy

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The Fermi energy E_F of an electron gas at temperature T is related to the value E_{Fo} at T=0 by the equation

$$E_F = E_{F_0} \left[1 - \frac{\pi^2}{12} \left(\frac{kT}{E_{F_0}} \right)^2 \right]$$

It is seen that as T is increased E_F decreases. But the rate of decrease with temperature is very small over a large range of temperature. The temperature at which E_F becomes zero is very large. For free electrons in metallic copper this temperature is of the order of 90 x 30^{3} K. So the temperature dependence of E_F over practical range of temperature be neglected and E_F at temperature T may be considered as constant equal to E_{F} .

Energy distribution curve

The Fermi –Dirac energy distribution law is

(1) Curve at T = 0. At T = 0, when $E < E_{F_0}$, then

$$e^{(E-E_{Fo})/kT} = e^{-\infty} = 0$$

From eq. (1)

n (E)dE =
$$4\pi V \left(\frac{2m}{h^2}\right)^{3/2} E^{1/2} dE$$

or n(E) = CE^{1/2} ------ (2)

where C is constant.

At T = 0, when $E = E_{F_0}$, then

$$e^{(E-E_{Fo})/kT} = e^{\infty} = \infty$$

Therefore under this condition n(E) = 0. Thus the curve representing eq. (2) is a parabola which ends abruptly at $E = E_{Fo}$.

(2) Curve at T>0:

At T>0, when $E << E_{Fo}$ the exponential term in the denominator of equ. (1) can be neglected in comparison with 1. So the region ($E << E_{Fo}$) is the same parabola. Then as E approaches E_{Fo} , the curves falls towards the axis of E and intersects the line $E=E_{Fo}$ at the point P. When $E>E_{Fo}$ the curve approaches the axis asymptotically showing that at a higher temperature a few electrons have energy greater than E_{Fo}

Fermi energy E_{Fo} for free electrons in a metal

The total number N of the free electrons in a metal of volume V is given by

$$N = \int_{0}^{\infty} n(E) dE$$

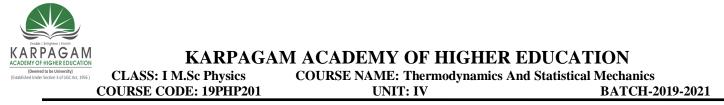
= $\int_{0}^{\infty} f(E) g(E) dE$
= $\int_{0}^{E_{F_{0}}} f(E)g(E) dE + \int_{E_{F_{0}}}^{0} f(E)g(E) dE$ ------ (1)

At T = 0, when $E \le E_{F_0}$, then f (E) =1 and at T=0, when $E \ge E_{F_0}$ then f(E) = 0. Hence in equ. (1) the second integral is zero and in the first integral f (E) = 1.

$$N = \int_{0}^{E_{F_{0}}} g(E) dE$$

= $4\pi V (2m/h^{2})^{3/2} \int_{0}^{E_{F_{0}}} E^{1/2} dE$
= $4\pi V (2m/h^{2})^{3/2} 2/3 E_{F_{0}}^{3/2}$
= $\frac{8\pi}{3} V \left(\frac{2mE_{F_{0}}}{h^{2}}\right)^{3/2}$
From this equation,
 $\left(\frac{2mE_{F_{0}}}{h^{2}}\right)^{3/2} = 3N / 8\pi V$
 $E_{F_{0}} = \frac{h^{2}}{2m} \left(\frac{3N}{8\pi V}\right)^{2/3}$ ------ (2)

When n = N/V = no.of free electrons per unit volume, i.e., the free electron density.



The values of E_{Fo} calculated from eq. (3) for a number of metals are of the order of several electron volts. This fact is a very important difference between classical statistics and Fermi-Dirac statistics. According to classical statistics all electron in a metal at absolute zero would have zero energy.

Fermi Temperature

The Fermi temperature T_h is defined as the ratio of the Fermi energy $E_{Fo}\,at$ absolute zero to Boltzmann's constant k. Thus

This equation shows that the degeneracy condition that A>1 is equivalent to $T_F>1$

The value of T_F for the free electrons in metal is very large. For the free electrons in copper $T_F 8.15 \times 10^4$ K. This value is much higher than room temperature so that the free electrons gas in copper is highly degenerate.

Interpretation of the Fermi Temperature

The temp. T of an ideal gas whose mean molecular K E is equal to the mean energy of the free electron in a metal at absolute zero is given by

 $3/2 \text{ kT} = U_0 = 3/5 \text{ E}_{Fo}$

Or
$$T = 2/5$$
 (E_{Fo}/k) = $2/5T_F$

It means that if free electron gas is considered to obey the classical statistics a piece of copper would have to be heated to a temp. 2/5 times the Fermi temperature for the metal is

2/5x8.15x10⁴=4.07x10⁴k

Thermionic Emission

According to the Fermi-Dirac statistics the free electrons in the highest energy level in a metal at absolute zero have the Fermi energy E_{Fo} . But they are not emitted spontaneously from the metal because of attractive forces of other charges at the surface. Therefore, to enable an electron having maximum energy E_{Fo} in the metal at zero K to escape from the surface to vacuum a certain minimum amount of energy Φ_0 must be important to it. This energy is called the work function of the metal at 0K. At a higher temperature these energies are denoted by E_F and Φ .

On heating a metal to a high temperature T a free electron with energy E_{Fo} may acquire an additional kinetic energy equal to the work function Φ of the metal. Then its total kinetic energy inside the metal is ($E_{F0} + \Phi$). The electron with this K.E inside the metal will just escape into vacuum from the metal surface, and will have 0K.E. on emergence. Hence this energy is the P.E. of the electron at rest outside the metal. This P.E. is called surface barrier of the metal. It is denoted by E_{S} . Thus $E_{S} = E_{Fo} + \Phi$.

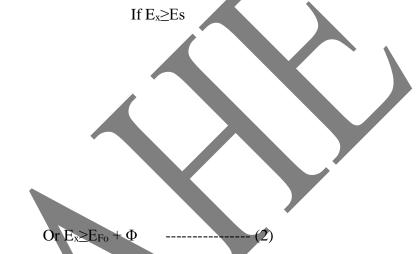
The Fermi-Dirac shows that if a metal is heated to a high temperature only those electron in the shaded portion of the curve will have K.E. > than Es and hence they can escape from the



metal surface. Thus when a metal is heated to a high temperature T a free electron will occur total K.E. > surface barrier potential energy Es will be emitted from the metal. This phenomenon is called thermionic emission.

Richardson – Dushmann Equation

Let E_x be the total K.E of an electron at temperature T in the positive x-direction. An electron having x-components of its velocity in the range between v_x and $v_x + dv_x$ will escape from the metal surface in the positive x-direction if $\frac{1}{2}$ mv_x² ≥Es.



The number of electrons per unit volume with the x-component of their velocity in the range v_x and v_x+dv_x is $\frac{n(v_x)dv_x}{v_x}$

The number of such electrons escaping from the metal surface per unit area per second is

$$v_x n(v_x) dv_x$$

The current density dJ due to these electrons that is the current per unit area i.e., the charge passing normally through unit area per second is given by

$$dJ = \frac{q v_{\chi} n (v_{\chi}) dv_{\chi}}{V} \quad -----(3)$$

where q – electronic charge. The total current density J is given by

The function $n(v_x)dv_x$ is given by Fermi-Dirac law of distribution of velocity in the xdirection

$$n(v_x)dv_x = V\left(\frac{4\pi m^2 kT}{h^3}\right)e^{E_{F_0/kT}} \cdot e^{\frac{-mv_x^2}{2kT}}dv_x$$



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- (7)

---- (8)

-- (9)

In this equation, we sub. $\frac{1}{2}$ mv_x² = E_x

$$m(v_x)dv_x = dE_x$$

$$dv_x = 1/mv_x dE_x$$
Hence, $n(v_x)dv_x = V\left(\frac{4\pi m^2 kT}{h^3}\right)e^{E_{F_0/kT}} \cdot e^{\frac{-E}{kT}} (1/mv_x)dE_x$

$$v_x n(v_x)dv_x = V\left(\frac{4\pi mkT}{h^3}\right)e^{\frac{E_{F_0}}{kT}} \cdot e^{\frac{-E}{kT}} dE_x$$

Or

Sub. this equation in equ. (4)

$$\mathbf{J} = \mathbf{q} \left(\frac{4\pi m kT}{h^3}\right) e^{\frac{E_{F_0}}{kT}} \int_{E_s}^{\infty} e^{\frac{-Ex}{kT}} \mathbf{d} \mathbf{E}_{\mathbf{x}} \quad ----- (6)$$

The value of the integral is

$$\int_{Es}^{\infty} e^{\frac{-Ex}{kT}} dE_x = \left[-kTe^{\frac{-Ex}{kT}}\right]_{Es}^{\infty}$$
$$= kTe^{\frac{-Es}{kT}}$$

$$= kTe^{-(EFo + d)}$$

)/kT

Sub. this value in eq. (6) and simplifying

$$J = \left(\frac{4\pi k^2 m q}{T}\right) T^2 e^{-\Phi/kT}$$

which is written as $J = A_0 T 2^2 e^{-\Phi/kT}$

Where $A_0 = \left(\frac{4\pi k^2 mq}{r^2}\right)$ = universal constant = 1.204 x 10⁶ A/m²K²

Eq.(9) is known as Richardson's or Dushman's or the Richardson- Dushmann equation. It was first derived by O.W. Richardson in 1901 and later the theory of its derivation were perfected by S.Dushman in 1923.

The eq. is based on the assumption that all the electrons having energy equal to or greater than the surface potential barrier or emitted but quantum mechanical theory shows that an electron having energy $E_x \ge E_s$ may not escape from the metal surface it may be reflected back, and that the probability of escape for such an electron is (1-r) where r is the reflection coefficient which is a function of (v_x) . But in thermionic emission velocity range of the emitted electron is not large, in this process r can be considered as constant over the small velocity range.

> $J=A_0(1-r)T^2e^{-\Phi/kT}$ -----(10)



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Comparison of M-B, B-E and F-D Statistics

Quantity	M-B	B-E	F-D
Particles	Distinguishable	Indistinguishable called	Indistinguishable called
		bosons	Fermions
Spin	-	0,1,2,	1/2, 3/2, 5/2,
Wave	-	Symmetric under	Antisymmetric under
function		interchange of two bosons	interchange of two bosons
Number of	No upper limit	Bosons don't obey Pauli	Fermions obey Pauli
particles per		exclusion principle: No	exclusion principle: Max.
energy state		upper limit to the no. of	of one particles per
		particles per quantum	quantum state.
		state.	
Distribution	$\frac{N}{V} \left(\frac{h^2}{2\pi m kT}\right)^{3/2} e^{-E/kT}$	1	
function f(E)	V \2mmint	$e^{\alpha}e^{\frac{E}{kT}}-1$	$e^{\frac{(E-Ef)}{kT}} + 1$



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Possible Questions

- 1. Give three differences between M-B, F-D and B-E statistics.
- 2. Explain Bose temperature and Bose-Einstein condensation
- 3. Which kind of particles is called as fermions?
- 4. What is called Bose- Einstein condensation?
- 5. Derive Richardson Dushmann equation.
- 6. Define Fermi energy.
- 7. Define the term Bose temperature.
- 8. What are called fermions?
- 9. Obtain Planck's law of radiation from Bose- Einstein distribution law.
- 10. Explain energy distribution curve of FD statistics.

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(For the candidates admitted from 2018 onwards)

DEPARTMENT OF PHYSICS

UNIT IV (Objective Type/Multiple choice Questions each Questions carry one Mark)

THERMODYNAMICS AND STATISTICAL MECHANICS

IAMI A(Omine Examination)									
S.No.	QUESTIONS	OPTION 1	OPTION 2	OPTION 3	OPTION 4	ANSWER			
1	B.E distribution law is used to derive of radiation	Planck's law	Weiss law	Widemann Franz law	Rayleigh's law	Planck's law			
2	Wave function of the system of identical Bosons is	asymmetric	linear	non-linear	symmetric	symmetric			
3	The variable W in an equilibrium stands for	minimum probability distribution	probability distribution	maximum probability distribution	constant probability distribution	maximum probability distribution			
4	Which of the following obey Pauli exclusion principle?	M.B. statistics	B.E. statistics	F.D. statistics	Einstein's equation	F.D. statistics			
5	Which of the following do not obey Pauli exclusion principle?	M.B. statistics	B.E. statistics	F.D. statistics	Einstein's equation	B.E. statistics			
6	In B.E. distribution, the constant e^{α} must be	greater than 1	smaller than 1	equal to 1	zero	greater than 1			
7	The molecule of an ordinary gas have spin angular momentum equal to an integral multiple of	h	ħ/2π	ĥ	h/2π	ħ			
8	The molecules obey B.E. statistics are	photons	phonons	fermions	bosons	bosons			
9	is the energy distribution	f(E)	g(E)	n(E)	f	f(E)			

PART –A(Online Examination)



	function.					
10	is the number of quantum states	f(E)	g(E)	n(E)	f	g(E)
11		postive infinity	negative infinitive	1	0	postive infinity
12	In B.E. energy distribution, if $e^{\alpha} < 1$, for E=0, then n(0) dE =	postive infinity	negative infinitive	1	0	negative infinitive
13	For an ideal B.E. distribution the degeneracy parameter A cannot be	greater than 1	smaller than 1	equal to 1	zero	greater than 1
14	For all known B.E. gases, T _B is very	high	low	small	0	low
15	The value of Reimann Zeta function is	6.212	1.612	2.126	2.612	2.612
16	A gaseous phase consisting of Ne molecules distributed among the energy states than the ground state.	higher	lower	equal to	very smaller	higher
17	The transition of liquid $_2\text{He}^4$ -I to superfluid liquid $_2\text{He}^4$ -II is observed at	8.12 K	2.18 K	1.82 K	1.28 K	2.18 K
18	For an ideal B.E. gas the condensation temperature is T_B . Find the temperature at which the number of molecules in the zero energy state (E=0) is 7/8 times the total number of molecules in the gas.	T=T _B	T=4T _B	T=1/4 T _B	T=4/3 T _B	T=1/4 T _B
19	E.Fermi developed the statistics for	photons	bosons	phonons	electrons	electrons
20	At T=0, when $E < E_{F0}$, the Fermi energy is given by	one	infinity	negative infity	zero	one
21	At T=0, when $E < E_{F0}$, the Fermi energy is given by	one	infinity	negative infity	zero	zero
22	For free electrons in metallic copper the temperature is of the order of	$90 \ge 10^3 \text{ K}$	90 x 10 ⁻³ K	$90 \times 30^3 \text{ K}$	90 x 30 ⁻³ K	$90 \times 30^3 \text{ K}$

23	For free electrons in copper, T _F =	$8.15 \times 10^2 K$	8.15 x 10 ⁻² K	8.15 x 10 ⁻⁴ K	8.15 x 10 ⁴ K	8.15 x 10 ⁴ K
24	The value of T_F free electrons in a metal is	very large	very small	zero	infinity	very large
25	The free electron gas in copper is	highly degenerate	degenerate	weekly degenerate	non-degenerate	highly degenerate
26	Surface potential barrier energy of the metal is denoted by	E _P	Es	S	S _E	Es
27	Richardson explain his theory in the year	1801	1701	1901	1921	1901
28	Spin value of bosons are in the order of	0,1,2,	0, 2,4,6,	¹ / ₂ , 3/2, 5/2,	1/3, 3/3, 5/3,	0,1,2,
29	S.Dushmann explained his theory in the year	1801	1701	1923	1921	1923
30	Spin value of fermions are in the order	0,1,2,	0, 2,4,6,	¹ / ₂ , 3/2, 5/2,	1/3, 3/3, 5/3,	1/2, 3/2, 5/2,
31	The value of b is given by	3 KT	KT	1/KT	4KT	1/KT
32	In B.E statistics the particles are identical and indistinguishable. These particles are called as	Bosons	fermions	leptons	baryons	Bosons
33	Particles with half-integral spin are called as	Bosons	fermions	leptons	baryons	fermions
34	Fermions obeyprinciple	Heisenberg	Le-chatlier	Pauli	Haber	Pauli
35		A<=1	A<<1	A>>1	A>>1	A<<1
26	In B.E statistics the particles are identical and indistinguishable. These	Deserve	6i	lantan	h	Deserve
36	particles	Bosons	fermions	leptons	baryons	Bosons
37	The Bosons has	spin 1	zero or half-	zero or whole	zero	zero or half-

			integral spin	number		integral spin
38	The examples for Bosons	photons	electrons	neutrons	protons	photons
	Particles with half-integral spin are					
39	called as	bosons	Fermions	leptons	electrons	leptons
40	The examples for Fermions	Photons	phonons	electrons	antiparitcles	electrons
41	The spin of the photon is	0	1	2	1/2	1
	In F.D statistics the particles are					
	identical and indistinguishable. These					
42	particles	Fermions	bosons	photons	kryptons	Fermions
			$\{1/(ea + bE)$	$\{(ea + bE) -$	$\{1/(ea + bE)$	$\{1/(ea + bE) -$
43		$\{1/(ea + bE)\}$	+1}	1}	-1}	1}
44	Fermi energy Ef =	- aKT	aKT	−1/ aKT	1/ aKT	- aKT
	Fermi-Dirac distribution function		${1/(ea + bE)}$ -	$\{1(ea + bE) -$	$\{-1/(ea + bE)$	$\{1/(ea + bE) +$
45	FD(E) =	$\{1/(ea + bE) + 1\}$	1}	1}	+1}	1}
	In terms of Fermi energy F.D		{1/(eE - EF)	$\{1/(eE + EF)\}$	$\{1/(eE + EF)$	$\{1/(eE + EF) +$
46	distribution function is fFD(E) =	$\{-1/(eE + EF) + 1\}$	+1}	+1}	- 1 }	1}
47	When $T = 0$ and $E < Ef$, then $fFD(E) =$	0	1	1/2	3	1
		N/V (h2 / 2pmkt	N/V (h2 /	N/V (h2 /	N/V (h2 /	N/V (h2 / 2pmkt
48	The degeneracy parameter e-a =)1/2	2pmkt)3/4	2pmkt)	2pmkt)3)3
	The maximum value of degeneracy					
49	parameter in B.E statistics is	One	two	three	Four	One
	B.E statistics is used to find the		frequency		mass	energy
50	among identical	energy distribution	distribution	both a and b	distribution	distribution
	According to B.E distribution law the	$g(E)dE/{(ea + bE) +}$	$g(E)dE/{(ea +$	$g(E)dE/{(ea +$	$g(E)dE/{(ea +$	$g(E)dE/{(ea +$
51	number of Bosons having energies	1}	bE) – 1}	bE) – 1}	bE) - 1 $1/2$	$bE) - 1 \} 1/2$
	According to Plank's law of radiation,					
52		Constant	same	not distinct	Vary	Constant
	For an isolated system the total energy					
	in a B.E. distribution law is used to					
53	derive of radiation	A<=1	A<<1	A>>1	A=1	A<<1
_	Wave function of the system of	.			Anti	
54	identical Bosons is	Unsymmetric	linear	symmetric	symmetric	symmetric

	The molecule of an ideal B.E gas in two					
55	phases at	T = TB	T > TB	T < TB	$T \leq TB$	T < TB
	Theconsists of no					
56	molecules occupying ground state	condensed phase	liquid phase	gaseous phase	inert gas	gaseous phase
	of same energy can have					
57	two different directions of polarization	Planck's law	Weiss law	Franz law	Newtons law	Planck's law
	Atthe molecule just reach		Kelvin	Neel	Curie	
58	the zero energy state	Bohr temperature	temperature	temperature	temperature	Bohr temperature



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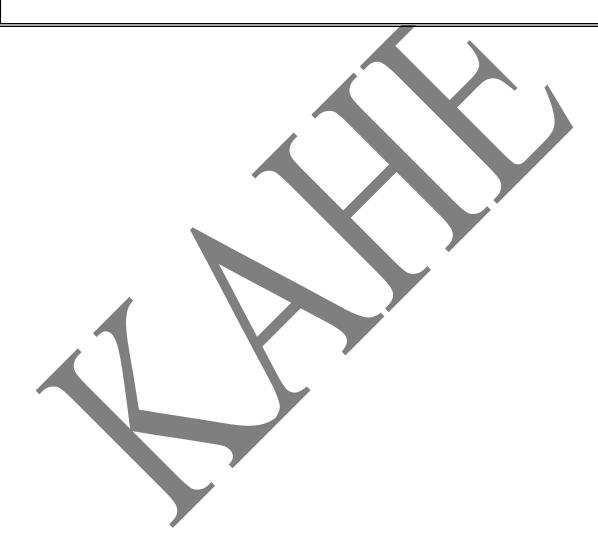
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Applications of Quantum Statistical Mechanics: Ideal Bose gas: Photons – Black body and

Planck radiation – Photons – Specific heat of solids – Liquid Helium.

Ideal Fermi gas: Properties – Degeneracy – Electron gas – Pauli paramagnetism

Ferromagnetism: Ising and Heisenberg models.





EMISSIVE POWER

The emissive power e_{λ} of a body at a given temperature for radiant energy of wavelength λ is defined as the energy emitted per unit area per seconds per unit range of wavelength $\frac{b}{n} \lambda \& \lambda + d \lambda$ at the given temperature.

unit \rightarrow 1 Walt per square meter per Angstrom.

ABSORPTIVE POWER (a_λ)

The absorptive power a_{λ} of a body at a given temperature for radiant energy of wavelength λ is the ratio of radiant energy of wavelength λ is absorbed per unit area per second, by the body at the given temperature to the radiant energy of the same wavelength incident on unit area per second at the temperature.

For perfectly black body $a_{\lambda}=1$.

Kirchoff"s law of Radiation

The law states that the ratio of the emissive power e_{λ} of a body for any wavelength at a given temperature to its absorptive power a_{λ} for the wavelength at the temperature is constant and is equal to the emissive power E_{λ} of a perfectly black body at that temperature.

=Eλ

Black Body Radiation

A body which completely absorbs radiation of all wavelength incident on it is called *perfectly black body*. Since a good absorber of radiation is also a good emitter of radiation, a perfectly black body is the best possible emitter at any given temperature .The radiation emitted by such kind of body is called black body radiation, or full radiation or temperature radiation.

A perfectly black body is an ideal conception. There is no known surface which can be regarded as perfectly black. Lamp black or platinum black is the nearest approach to the perfectly black body. Lamp black can absorb about 96% of the radiant energy incident on it, &platinum black absorbs about 98%. In practice almost perfectly black body consist of a double-walled hollow metal sphere. The sphere has a small hole O. There is conical projection P Prepared by Dr.S.Sharmila, Assoc Prof, Department of Physics, KAHE 2 | 10



opposite the hole. The inner surface of the sphere is coated with lamp black .The space between the wall is evacuated to prevent loss of heat by conduction or convention .when any radiant energy enters the space through the hole, it suffers multiple reflections. At each reflection about 96% of the incident radiant energy is absorbed .hence after few reflections all the radiant energy is absorbed by the sphere. The function of the conical projection is to prevent direct reflection of the radiant energy from the surface opposite to the hole. The hole O in the sphere acts as a perfectly black body because it absorbs all the radiant energy incident on it. When the sphere is heated black body radiation is emitted from the hole.

Liquid Helium

As an application of B-E. statistics, the qualitative nature of the superfluid transition of liquid helium at 2.2 K were investigated. Ordinary helium consists almost entirely of neutral atoms of the isotope ${}_{2}\text{He}^{4}$. As the total angular momentum of this atom is 0.

Helium exhibits peculiar properties at low temperatures.

- Helium gas at atmospheric pressure condenses at 4.3 K into a liquid of very low density about 0.124 gm/cm³
- (ii) Further cooling about 0.82 K doesnot freeze it and its believed that it remains all the way down to absolute 0. The solid state of helium does not form unless it is subjected to an external pressure of atleast 23 atm.
- (iii) For He¹ in liquid phase there is another phase transition called λ transition which divides the liquid state into two phases He I and II. K. Onnes while liquefying helium noted that about 2.2 K, density appeared to pass through abrupt maximum and then decreasing slightly. Investigations also revealed that critical temperature at 2.186 K. It represents a transition to a new state of matter known as liquid He II.
 - a) Heat conductivity is very large in the order of 3.10^6 times greater.
 - b) Co-efficient of velocity gradually diminishes as the temperature is lowered and appears to be approaching 0 at absolute 0 temp.
 - c) Specific heat measurements by Kessom show that specific heat curve is discontinuous at 2.186 K. The shape of the specific heat curve resembles the shape of the letter λ and this peculiar transition is called λ transition and the



discontinuity temp. 2.186 K is called λ point. Kesson concluded that transition He I \rightarrow He II at T_{λ} is second order transition. The transition temp. decreases as the pressure is increased.

Degeneracy

(i) Weak degeneracy: At $T > T_F$ (i.e., at intermediate temperatures) the Fermi gas is said to be slightly degenerate. In this case $kT > \in_F (0)$, then \in_F is negative or α is positive and A < 1.

$$n = g_{s} \cdot V/h^{3} (2\pi mkT)^{3/2} A$$

$$E = 3/2 g_{s} \cdot V/h^{3} (2\pi mkT)^{3/2} kT.A$$

$$E/n = 3/2 kT \text{ or } E = 3/2 nkT \qquad ------(6)$$

which is well known relation for a perfect gas in classical statistics.

A comparison of equations (5) and (6) shows that ideal Fermi-Dirac gas deviates from perfect gas behaviour and this derivation is called degeneracy. It is obvious that degeneracy is the function of A. Greater is the value of A, more marked will be the degeneracy. Hence for A < 1 or $T > T_F$, the Fermi gas slightly degenerate.

Case (ii) strong degeneracy. When α is large and negative A=e^{- α} >>1. As degeneration increases will increases of A, therefore in this case degeneracy becomes more prominent. Further to the first approximation; from en. (3)

$$A \approx 1/g_s \cdot n/V h^3/(2\pi m kT)^{3/2}$$

This eqn. shows that the gas will be strongly degenerate at low temperature and high particles densities n/v. The evaluation of integrals $f_1(\alpha)$ and $f_2(\alpha)$ under these conditions is complicated.

This case of strong degenerate at low temperature ranges:

(a) At absolute zero i.e when T=0

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(b) When T is above absolute zero, but A >> 1.

Case (a) At absolute zero i.e. when T=0. When $T \rightarrow 0 A \rightarrow 0$. In this case the Fermi-dirac gas is completely degenerate.

At T=0,

$$f(\epsilon) = \frac{1}{\frac{1}{A}e^{\epsilon/kT} + 1} = \frac{1}{e^{\frac{(\epsilon - \epsilon_F)}{kT}} + 1} = 1 \text{ for } 0 \le \epsilon \le \epsilon_F(0)$$

=0 for $\epsilon > \epsilon_{\rm F}(0)$

where $\epsilon_{\rm F}(0)$ is given by eqn. (18)

Now the total internal energy of perfect Fermi- dirac gas at T=0 i.e. zero point energy of Fermi gas is

$$E_0 = 3nh^2/10m [3n/4\pi Vg_s]^{2/3} = 3/5 n \epsilon_F(0) \qquad (8)$$

Now the pressure at T=0 is given by

 $P_0 = 1/3 E_0/V$ = 1/5nh²/V_m (3n/4\pi g_sV)^{2/3} ------ (9)

Form equations (8) and (9) it is obvious that a strongly degenerate Fermi-dirac gas possesses energy and exerts a pressure even at 0K, quite unlike a Bose Einstien and classical gases where the energy and pressure at absolute zero are zero.

Case (b) At temperature above absolute zero :but A>>1or T<<Tf

In this case the Fermi –gas is strongly degenerate at low temperature and ϵ is still positive. From equation (17) the number of particles lying in the energy range between ϵ and ϵ +d ϵ is given by

dn (
$$\epsilon$$
) =f(ϵ)d(ϵ) = 3/2 n/[ϵ_F (0)]^{3/2}. $\epsilon^{1/2}$ d $\epsilon/e^{(\epsilon-\epsilon F)/kT} + 1$ -----(10)

Therefore, the total number of particles is



and the total integral energy is

To solve the integrals in equations (31) and (32), let us consider the general integral of the type

Where $\emptyset(\epsilon)$ is a simple function of such that $\emptyset() = 0$ when =0.

The integral of eqn. (33) can be expanded using the method of taylor's series expansion,

----- (15)

Where \emptyset', \emptyset''' etc. denote the first, third etc. differentials of the function \emptyset .

Now for $\Phi(\epsilon) = \epsilon^{1/2}$

$$\epsilon_{\rm F} / \epsilon_{\rm F} (0) = [1 + 1/8(\pi k T / \epsilon_{\rm F})^2 + 7/640 (\pi k T / \epsilon_{\rm F})^4 +]^{-2/3}$$
 (16)

Remembering that kT < < f, we can take into account only the first two terms in the bracketed expression and write

$$\epsilon_{\rm F}/\epsilon_{\rm F}(0) = [1 + 1/8(\pi k T/\epsilon F)^2]^2 \approx 1 - 1/12 (\pi k T/\epsilon F)^2$$
 ------(17)

This gives

$$1/\epsilon_{\rm F}^2 \approx 1/[\epsilon_{\rm F}(0)]^2 [1+1/6 (\pi k T/\epsilon F)^2]$$
 ------ (18)

Now make the crude approximation by putting f = f(0) in the second term of above expression

 $1/\epsilon_{\rm F}^2 \approx 1/[\epsilon_{\rm F}(0)]^2 [1+1/6 (\pi kT/\epsilon_{\rm F}(0)^2]$ ------(19)

Now using equation (19), equation (17) gives

$$\epsilon_{\rm F} \approx \epsilon_{\rm F} (0) \left[1 - 1/12 \left(\pi k T / \epsilon_{\rm F}(0) \right)^2 \right]$$
 ------ (20)

 $\mathbf{E} \approx \mathbf{n} \ \boldsymbol{\epsilon}_{\mathrm{F}} \left(0 \right) \left(\boldsymbol{\epsilon} \mathbf{F} \ / \boldsymbol{\epsilon}_{\mathrm{F}} \left(0 \right)^{5/2} \quad [1 + 5/8 \ (\pi \mathrm{kT} / \ \boldsymbol{\epsilon}_{\mathrm{F}})^2]$

Now using equations (39) and (40), we get

 $E \approx 3/5 \text{ n} \epsilon_F(0) [1+5/12(\pi kT/\epsilon_F(0))]$ ------(22)

The corresponding pressure is

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 $P = 2/3 E/V \cong 2/5 (n \epsilon_F (0)/V) [1+5/12 (\pi kT/\epsilon_F (0)^2]$ ------(23)

Equations (22) and (23) represent the approximate energy and pressure of a strongly degenerate gas.

UNIT: V

Electron Gas

A metal can be considered to be composed of a system of fixed positive nuclei and a no. of mobile electrons referred to us the electron gas.

To study the properties of an electron gas at low temperature in the region. For electrons s=1/2 so

$$g_z = 2s + 1 = 2$$

$$\epsilon_{\rm f} = {\rm h}^2/2{\rm m} \left({{3n}\over{4\pi~V.2}}\right)^{2/3}$$

= {\rm h}^2/8{\rm m} \left({{3n}\over{\pi~2}}\right)^{2/3}

$$E_a = 3/5 n \epsilon_f$$

In the limit $T \rightarrow 0$ Which means that in the limit every one of the states is occupied fully up to the energy level ϵ_f whereas all the states above this energy level are empty. For electrons $m = 9.1 \times 10^{-28}$ gm, and g = 2.

----- (2)

$$\frac{1}{D} = \frac{h^2}{2 x 9.1 x 10^{-28} kT} \left(\frac{3n}{8\pi V}\right)^{2/3}$$

A typical metal atomic weight 100 and density 10 so that the volume of gm. atom of 10cc. and the number of electrons assuming one free electron from atom is 6.02×10^{23} . Then

$$\frac{1}{D} = \frac{(6.62 \ x \ 10^{-27})^2}{2 \ x \ 9.1 \ x \ 10^{-28} \ x \ 1.38 \ x \ 10^{-18} \ x \ T} \left(\frac{3x \ 6.02 \ x \ 10^{23}}{8 \ x \ 3.14 \ x \ 10}\right)^{2/3}$$
$$= 10^{5/} \ 1.5T$$

Which means degeneracy is sufficiently high. It shows clearly that for electron gas, a classical statistics is not valid and can be applied only at temperature of the order of 10 5 K. Therefore at low and other ordinary working temperatures, it is necessary to use Fermi- Dirac statistics to study the electron gas in the metals. At low temperature electronic contribution to the specific heat of metals is given by the equ.

$$C_{v} = \frac{1}{2} \text{ nk } \pi^{2} (\text{kT}/\epsilon_{f})$$
$$D = (\text{kT}/\epsilon_{f})$$



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 $C_v = \frac{1}{2} nk \pi^2 D$

Using this value find the value of 1/D,

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$$C_v = \frac{1}{2} \text{ nk } \pi^2 \text{ x } 1.5 \text{ x } 10^{-5} \text{ x T}$$

Sub.
$$nk = R$$

$$= 1.987 \text{ cal deg}^{-1} \text{ mol}^{-1}$$

$$= 2 \text{ cal deg}^{-1} \text{ mol}^{-1}$$

$$\pi^2 = 10$$

$$C_v = \frac{1}{2} \ge 1.5 \ge 10^{-5} \ge 2 \ge 10 \ge T$$

$$= 1.5 \text{ x } 10^{-4} \text{ x T cal/gm. atom}$$

Pressure of the electron gas can be obtained by,

 $\mathbf{F}_{0} = \frac{1}{5} \frac{1}{V}$ $= \frac{2}{5V} \frac{h^{2}}{2m} \left(\frac{3n}{4\pi V gs}\right)^{2/3}$ $= \frac{nh^{2}}{20mV} \left(\frac{3n}{\pi V}\right)^{2/3} \text{ using } gs = 2$

For a metal of atomic weight 100 and density 10

 $P_{o} \sim 10^{5}$ atoms

Which means at normal temperature the pressure of gas is sufficiently high.

Ising and Heisenberg Model

Transition of non-ferromagnetic state into ferromagnetic state is called phase transition, in this transition, the state of the body changes continuously.

Consider a ferromagnetic substance, like iron and nickel. Without any external field being applied, some of the spins of the atoms become spontaneously polarized in the same direction, below curie temperature Tc. This create a macroscopic magnetic field. The spontaneous magnetization, created vanishes if temperature is greater than Tc, because thermal energy makes some of the aligned spin to flip over. Thus spins get oriented at random no net magnetic field is produced. As the curie temperature approached both sides of the specific heat of the metal approaches infinity. The transition from non-ferromagnetic to the ferromagnetic state, called the phase transition, is associated with some kind of change in the symmetry of the lattice; For example the ferromagnetism symmetry of the spins is involved. In Ising model the system considered is the array of N fixed points called lattice sites that from an n-dimensional periodic lattice (n=1,2,3). Associated with each lattice site is a spin variable, s_i , i = 1 to n, It is a number that either +1 or -1. There is no other variable. If $s_i = +1$ the ith state is said to have spin up and $s_i = -1$, it is said to have spin down. A given set of $\{s_i\}$ specifies a configuration of the whole system, whose energy is defined to be

Where the subscript I stands for Ising and the symbol $\langle i, j \rangle$ denotes a nearest-neighbour pari of spins. There is no distinction between $\langle i, j \rangle$ and $\langle j, i \rangle$ is the interaction energy μ H interaction is associated with the external magnetic field H. For spontaneous magnetization H= 0. \in_{ij} and H are given constants. Applied the model to the case is isotopic interaction so that all \in_{ij} have the same value \in . For energy

$$E_1\{s_i\} = -\in \sum_{\langle i,j \rangle} \quad si \ sj - \mu H \sum_i si$$

The case $\in > 0$ corresponds to ferromagnetism and the case $\in < 0$ to antiferromagnetism. In the former case neighbor spins tend to be parallel while in the latter case they tend to be antiparallel. In eq. (2) the sum over $\langle i, j \rangle$ contains N/2 terms where is the number of nearest neighbors of any given site. In the Ising model eq. 2, geometry of the lattice enters through and interaction energy \in_{ij}

Consider only these case $\in > 0$. The partition function is

$$Z = \sum_{S1} \sum_{S2} \sum_{SN} e^{-\beta E1 \{Si\}}$$

(2)



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Possible Questions

- 1. Define specific heat capacity at constant volume.
- 2. Explain Ising and Heisenberg model.
- 3. Discuss about specific heat of solids.
- 4. What are called photons?
- 5. State Dulong and Petit's law.
- 6. Give a note on liquid helium.
- 7. Define degeneracy.
- 8. Give a note on pauli paramagnetism.
- 9. What are the applications of quantum statistical mechanics?
- 10. State Kirchhoff's law.
- 11. Write a note on black body and Planck radiation law.

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DEPARTMENT OF PHYSICS

UNIT V (Objective Type/Multiple choice Questions each Questions carry one Mark)

THERMODYNAMICS AND STATISTICAL MECHANICS

	raki -A(Onnie Examination)									
S.No.	QUESTIONS	OPTION 1	OPTION 2	OPTION 3	OPTION 4	ANSWER				
1	In solid heat is transferred by	Conduction	convection	radiation	Irradiation	Conduction				
2	In liquid heat is transferred by	Conduction	convection	radiation	Irradiation	convection				
	Conduction and convection cannot take									
3	place in	solid	liquid	inert gas	empty space	empty space				
	does not require any									
4	material medium.	Conduction	convection	radiation	Irradiation	radiation				
5	Radiant energy is also called as	latent heat	radiant heat	entropy	enthalpy	radiant heat				
	The wavelength of infrared ranges from	7500Å to	750Å to	7500Å to	750Å to	7500Å to				
6		1000000Å	100000Å	10000Å	1000000Å	1000000Å				
7	travels with speed of light.	Conduction	convection	radiation	Irradiation	radiation				
	The nature of radiant energy is same as									
8	that of	sound	heat	light	electricity	light				
	The radiant energy emitted depends on									
9		temperature	material	volume	height	temperature				
	For a black body the emissive power is									
10	denoted by	Е	λ	λΕ	Ελ	Е				
11	Unit of emissive power is	1 W/m	1WmÅ	1WÅ	1W/m2Å	1W/m2Å				
12	For a perfectly black body $a\lambda =$	0	x	1	-∞-	0				

PART –A(**Online Examination**)



13	Absorptive power is represented by	a	λ	λα	aλ	aλ
14	Kirchoff's law of radiation =	$e\lambda/a\lambda = E\lambda$	$e\lambda/a\lambda = \lambda$	$a\lambda/e\lambda = E\lambda$	$e\lambda/E\lambda = a\lambda$	$e\lambda/a\lambda = E\lambda$
	black is nearest approach to					
15	a perfectly black body.	gold	platinum	diamond	silver	gold
	Which is considered as a perfect absorber					
16	I	Gray body	Black body	Real body	White body	Black body
	Which body that emits a constant					
17	emissivity regardless of the wavelength?	Gray body	Black body	Real body	White body	Gray body
	At same temperatures, the radiation					
	emitted by all real surfaces is the				Either less than or	
18	radiation emitted by a black body.	Less than	Greater than	Equal to	greater than	Less than
				It is		
				independent	T4 in 1,	T4 :- 1:41
	Which is NOT a characteristic of	It is high with	It is directly	with the surface	It is low with	It is low with
19		It is high with most nonmetals	proportional to	condition of the material	highly polished metals	highly polished metals
	emissivity?		temperature			Inetais
20	What is the emissivity of a black body?	0	1	0.5	0.25	<u> </u>
21	What is the absorptive of a black body?	0	1	0.5	0.25	0
				a insulating	a diamagnetic	
		A ferromagnetic		material	material becomes	A ferromagnetic
22	Alterna Carrie and int	material becomes	a ferrite becomes	becomes a	a paramagnetic	material becomes
22	Above Curie point	paramagnetic	an insulator	ferrite	material	paramagnetic
23	Which of the following is a paramagnetic material?	Palladium	Lead	Pure Iron	Bismuth	Lead
23	Which of the following is a ferromagnetic	Fallaululli	Leau	rute non	Disiliutii	
24	material?	Palladium	Lead	Iron	Bismuth	Iron
<u></u> _	All of the following materials are	1 unuurum	Loud	101	Distitutii	
25	Ferromagnet except	Nickel	Bismuth	Silicon	Mild steel	Mild steel
		electrical	electrical	electrical	electrical	electrical resistivity
		resistivity	resistivity	resistivity	resistivity	decreases and
	By adding silicon to ferromagnetic,	increases and also	decreases and	decreases and	increases and	magnetic
1	materials		also magnetic		1	permeability

		permeability	permeability	permeability	permeability	increases
		increases	decreases	increases	decreases.	
		mass of the	heat capacity/	mass of the		
	The specific heat capacity of a substance	substance \times heat	mass of the	substance/heat	mass of the	heat capacity/mass
27	is equal to	capacity	substance	capacity	substance	of the substance
28	Specific heat capacity of glass is	635 J kg ⁻¹ °C ⁻¹	670 J kg ⁻¹ °C ⁻¹	705 J kg ⁻¹ °C ⁻¹	740 J kg ⁻¹ °C ⁻¹	670 J kg ⁻¹ °C ⁻¹
				the amount of		
				heat required to	the amount of	
				change the	heat required to	
		the amount of heat	the amount of	phase of a	change the phase	the amount of heat
		required to raise	heat required to	substance from	of a substance	required to raise the
		the temperature of	raise the	solid to liquid	from liquid to gas	temperature of a 1
	The specific heat capacity of a substance	a 1 kg of a	temperature of a	without any	without any	kg of a substance
29	is equal to	substance by 1 K	substance by 1 K	chan	change	by 1 K
30	Specific heat capacity of mercury is	120 J kg ⁻¹ °C ⁻¹	140 J kg ⁻¹ °C ⁻¹	160 J kg ⁻¹ °C ⁻¹	180 J kg ⁻¹ °C ⁻¹	140 J kg ⁻¹ °C ⁻¹
	The amount of heat required to raise					
	temperature of a substance by 1°C is				specific heat	
31	called as:	work capcaity	heat capacity	energy capacity	capacity	heat capacity
		change in		nature of	height of	
32	Heat capcity does not depends on	temperature	mass of body	substance	substance	height of substance
33	Heat brings change	physical	chemical	reversible	periodic	chemical
	The amount of heat required to raise the				specific heat	specific heat
34	temperature of 1 kg by 1°C is called as:	work capcaity	heat capacity	energy capacity	capacity	capacity
35	SI unit of specific heat capacity is:	kg°C	j/kg°C	j/kg°	j/g°C	j/kg°C
	Which of the following has highest heat					
36	capacity?	water	air	soil	wood	water
	The temperature at which liquid changes					
37	into vapour is called as	Melting point	boiling point	expansion point	phase transition	boiling point
		without			without	
		themselves		themselves	themselves	without themselves
	In Conduction process the molecules of	moving from their		move from one	moving from one	moving from their
38	the solid pass the heat from one to another	positions	No movement	place to another	place to another	positions

	The process of transfer of heat in liquids &					
39	gases is called as	Conduction	Radiation	Convection	absorption	Convection
		solid are not free	molecules only			
	Solids are not heated by convection	to move from one	vibrate about		they are loosely	
40	because	place to another	fixed position	both A and B	packed	both A and B
	It is the process of heat transfer from a hot					
	body to a colder body without heating the					
41	space between the two is called as	Conduction	Radiation	Convection	absorption	Radiation
		does not require	require any	does not require		does not require
42	The transfer of heat by radiation	any medium.	medium	any space	require any space	any medium.
43	Heat of sun reach the earth by	Conduction	Radiation	Convection	absorption	Radiation
	A cold steel spoon is dipped in a cup of					
	hot milk. It transfers heat to its other end					
44	by the process of	Conduction	Radiation	Convection	absorption	Conduction
			heat is	heat is		
			transferred from	transferred from		
			the hotter end to	the colder end		
		particles of solids	the colder end of	to the hotter end		
45	Why conduction is only possible in solids	are closely packed	an object	of an object	both A and B	both A and B
	The water is poor conductors of heat so do					
46	not heated by	Conduction	Radiation	Convection	absorption	Conduction
	Which of the following are the examples				• 1 •	
47	of conductors?	plastic	Iron	wood	silicon	iron
40	Which of the following are the examples		т	1	•1•	1
48	of insulators ?	copper	Iron	wood	silicon	wood
10	Radiation is the transfer of heat by means		electromagnetic	1 4 * 1	1'	electromagnetic
49	of Materials as high hash as a set of the set	magnetic wave	waves	electrical wave	radio waves	waves
50	Materials which lack permanent magnetic	die weerst	former and en et			dia maanat
50	dipoles are called	dia magnet	ferro magnet	semi-magnet	para magnet	dia magnet
	Materials having a high dielectric constant, which is non-linear, are known		ferroelectric	super die		ferroelectric
51	· · · · ·	elastomers	materials	super die- electrics	hard die-electrics	materials
52	In ferromagnetic materials	the atomic	the atomic	the constituents	one of the	the atomic

		magnetic moments	magnetic	is iron only	constituent is iron	magnetic moments
		are antiparallel and	moments are			are parallel
		unequal	parallel			
	The temperature beyond which substances					
	lose their ferroelectric properties, is known		critical	inversion	conversion	
53	as	curie temperature	temperature	temperature	temperature	curie temperature
	What is the degeneracy of the rotational					
	energy level with $J = 4$ for a heteronuclear					
54	diatomic molecule?	1	2	3	9	9
	Which type of statistics is used to describe					
55	the electron contribution to specific heat?	MB statistics	BE statistics	FD statistics	Classical statistics	FD statistics